

From Organic Astrochemistry to Prebiotic Systems

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VI AstrobiON

Escola de Astrobiologia do Observatório Nacional, Rio de Janeiro, Brazil
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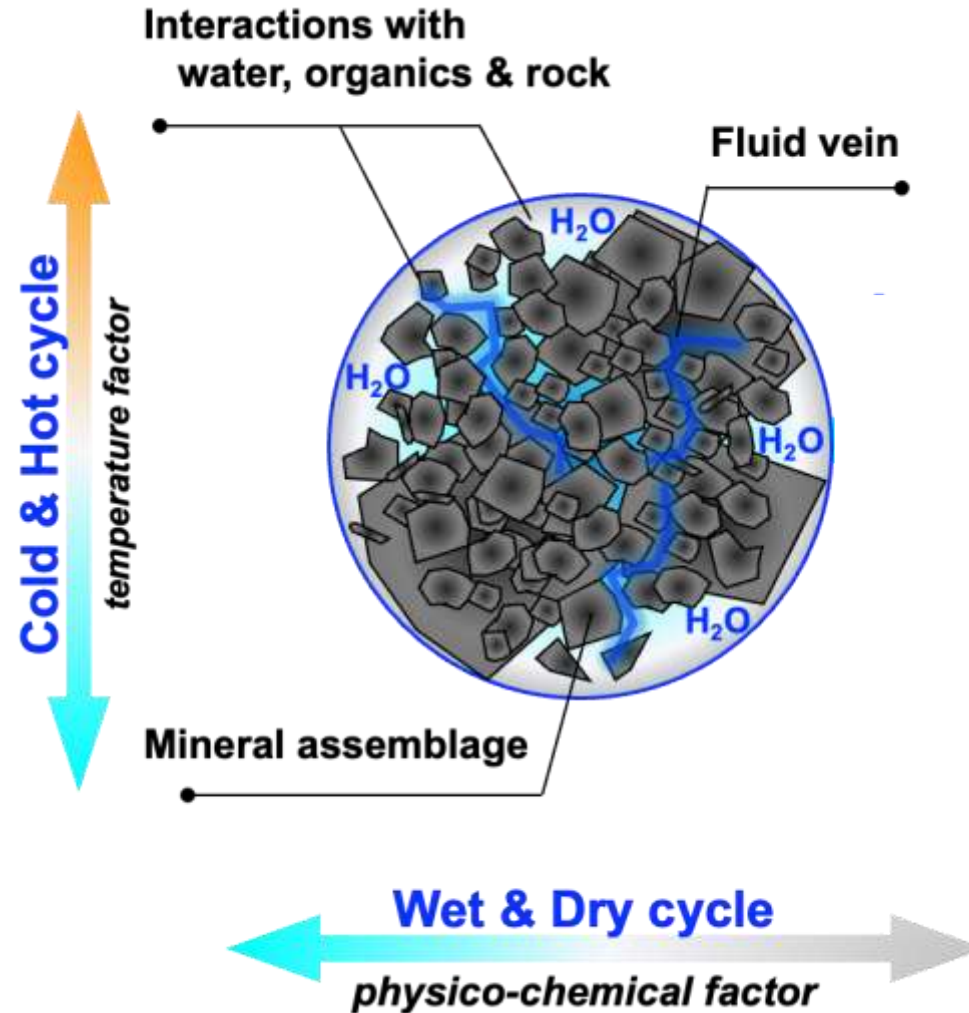
GSFC ASTROBIOLOGY ANALYTICAL LAB <http://science.gsfc.nasa.gov/691/analytical/>



What Drove Organic Synthesis in Meteorites?

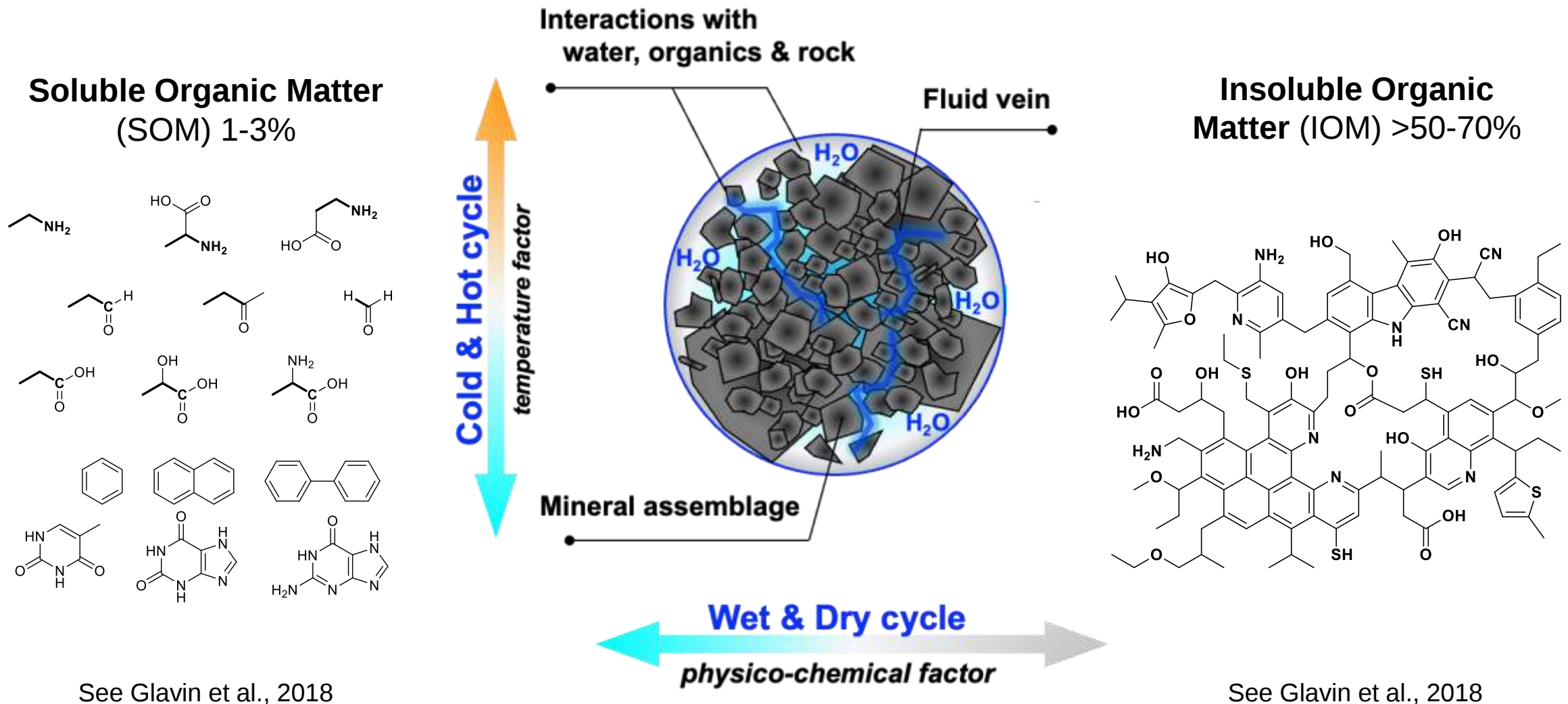
Parent-body aqueous and thermal alteration from melting of ices due to **radiogenic decay**, **tidal forces**, and **impacts** results in the formation of increasingly complex prebiotic molecules from:

- Strecker synthesis
- Condensation reactions
- Michael addition
- HCN polymerization
- Fischer-Tropsch-type reactions
- etc.



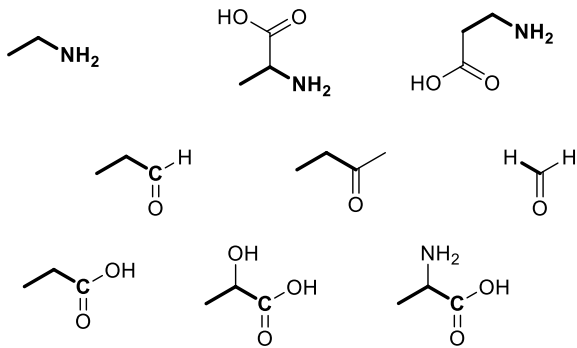
Modified from Takano et al., 2024

Organics in Carbonaceous Chondrites



Organics in Carbonaceous Chondrites

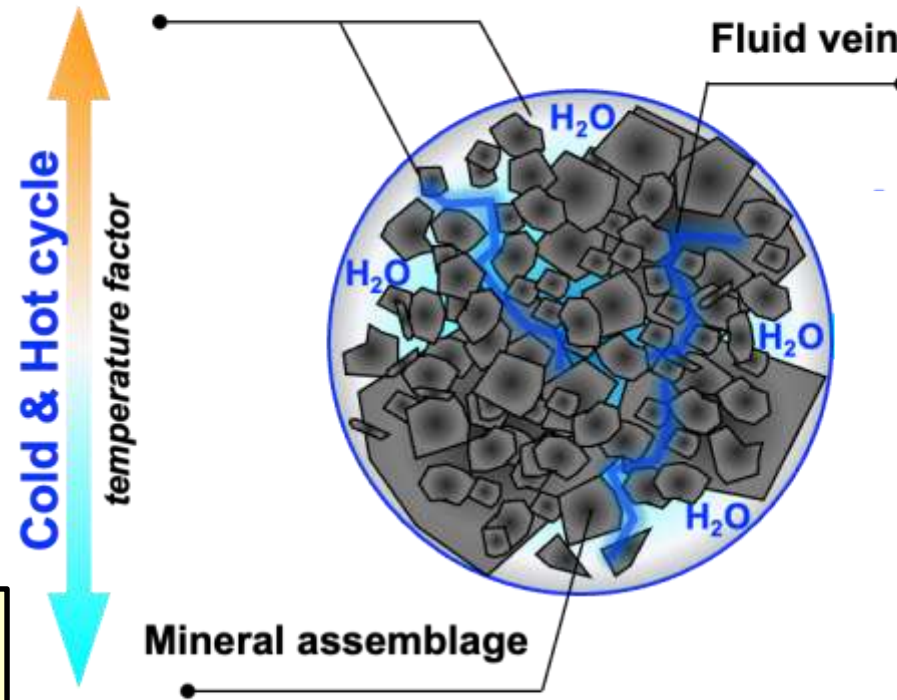
Soluble Organic Matter (SOM) 1-3%



- Feedstock availability
- Reaction pathways
- Interstellar vs parent-body
- Degree of alteration

See Glavin et al., 2018

Interactions with water, organics & rock

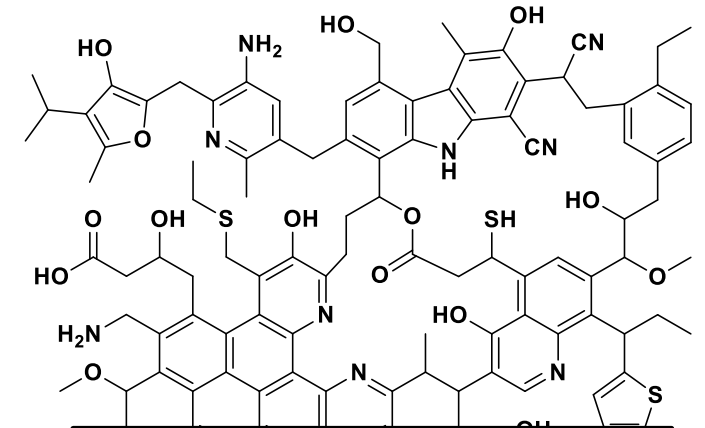


Cold & Hot cycle
temperature factor

Wet & Dry cycle

physico-chemical factor

Insoluble Organic Matter (IOM) >50-70%



- Degree of alteration
- Aqueous vs thermal

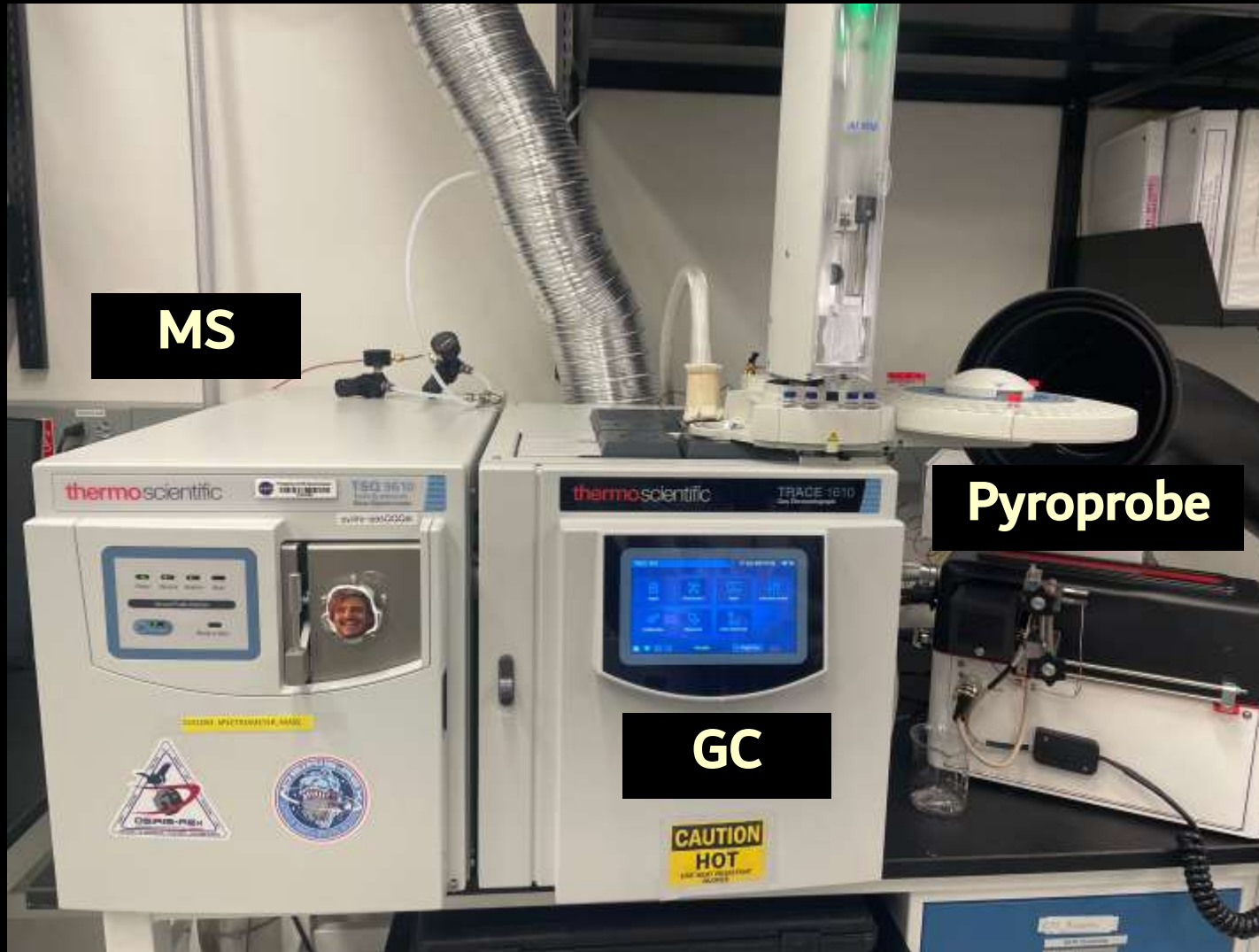
See Glavin et al., 2018

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Bulk Sample Analysis by pyrolysis-GC-MS

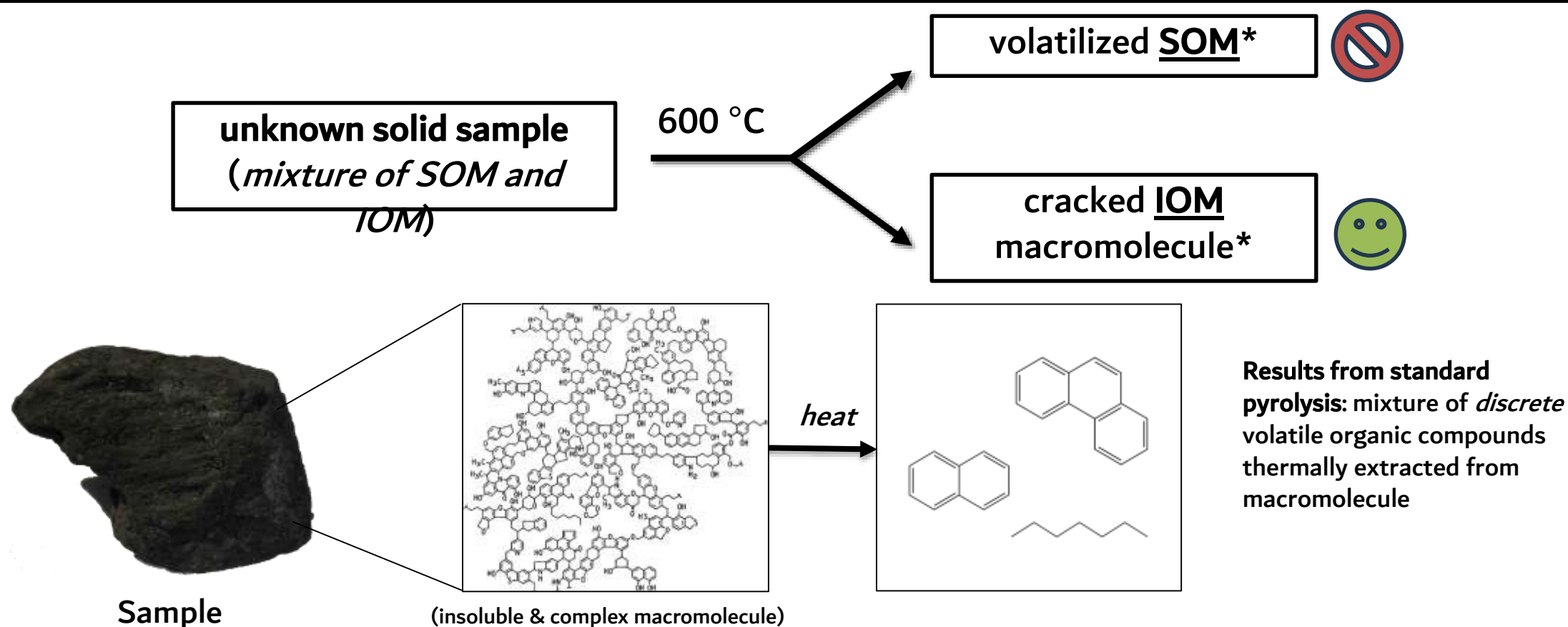
Pyrolysis-gas chromatography-mass spectrometry (py-GC-MS) characterizes complex organic mixtures

Bulk solid
samples
without
chemical
pretreatments
IOM + SOM



~1 mg
of sample

Standard Pyrolysis



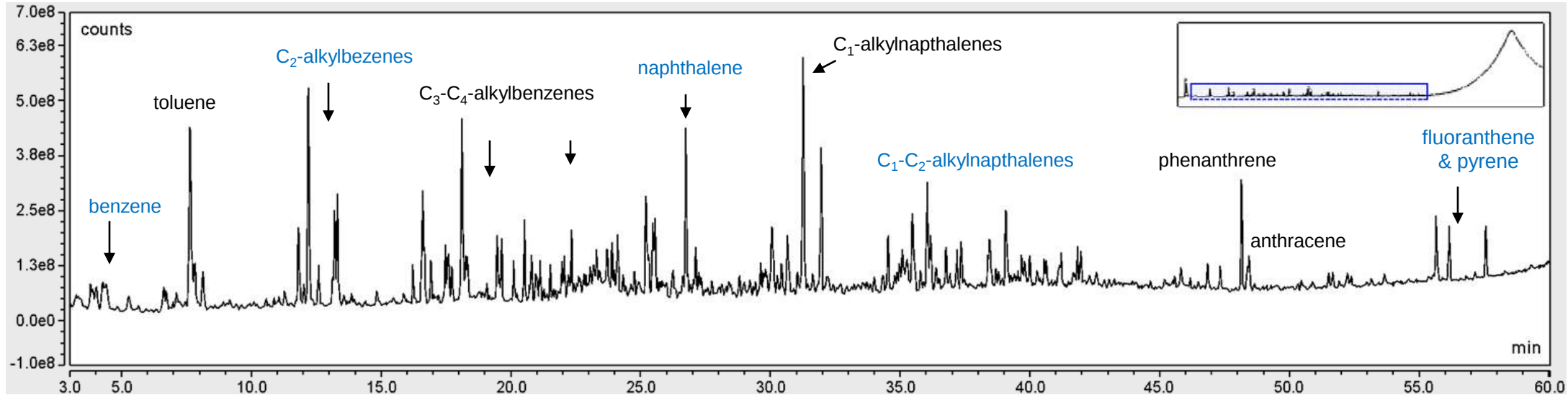
Reasoning: IOM wt.% is larger than soluble organics.

Pyrolysis signals represent the stable hydrocarbon fraction from the thermal breakdown of IOM.

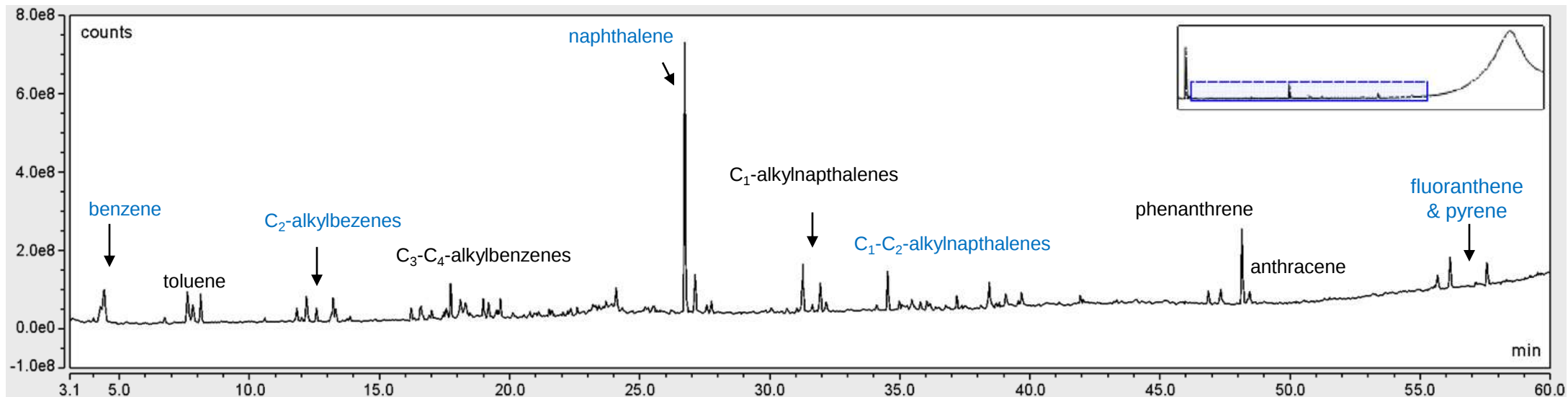
*Thermally unstable compounds may undergo dehydration, decarboxylation, etc. at the chosen pyrolysis temperature.

Pyrolysis Total Ion Chromatogram

**OREX-
800055-112**
angular



Murchison

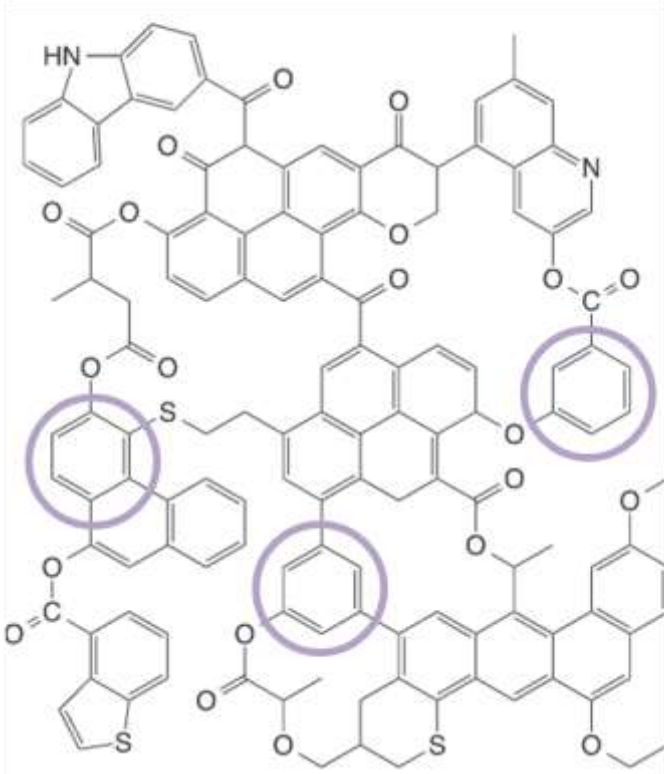


Comparison of PAH in Meteorites

Mojarro et al. (2025) *PNAS* e2512461122

IOM *conceptual* structure

(Glavin et al., 2013)



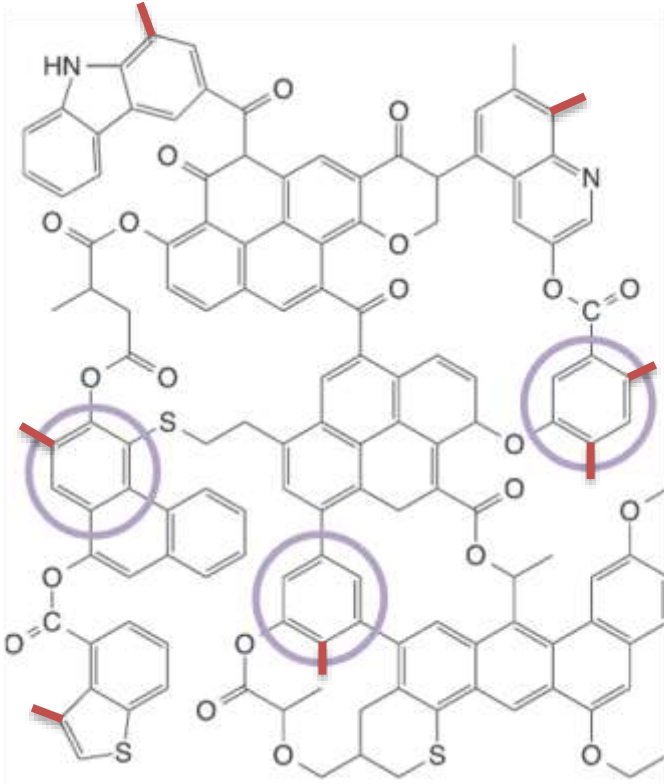
Hypothesis: Distribution of alkylated benzenes, heterocycle, & PAH isomers may reveal IOM alteration history

Comparison of PAH in Meteorites

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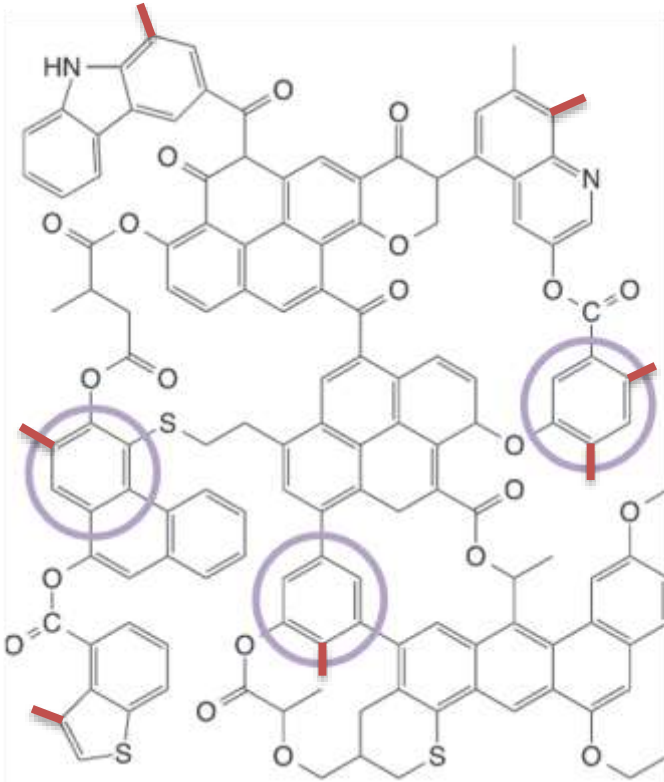
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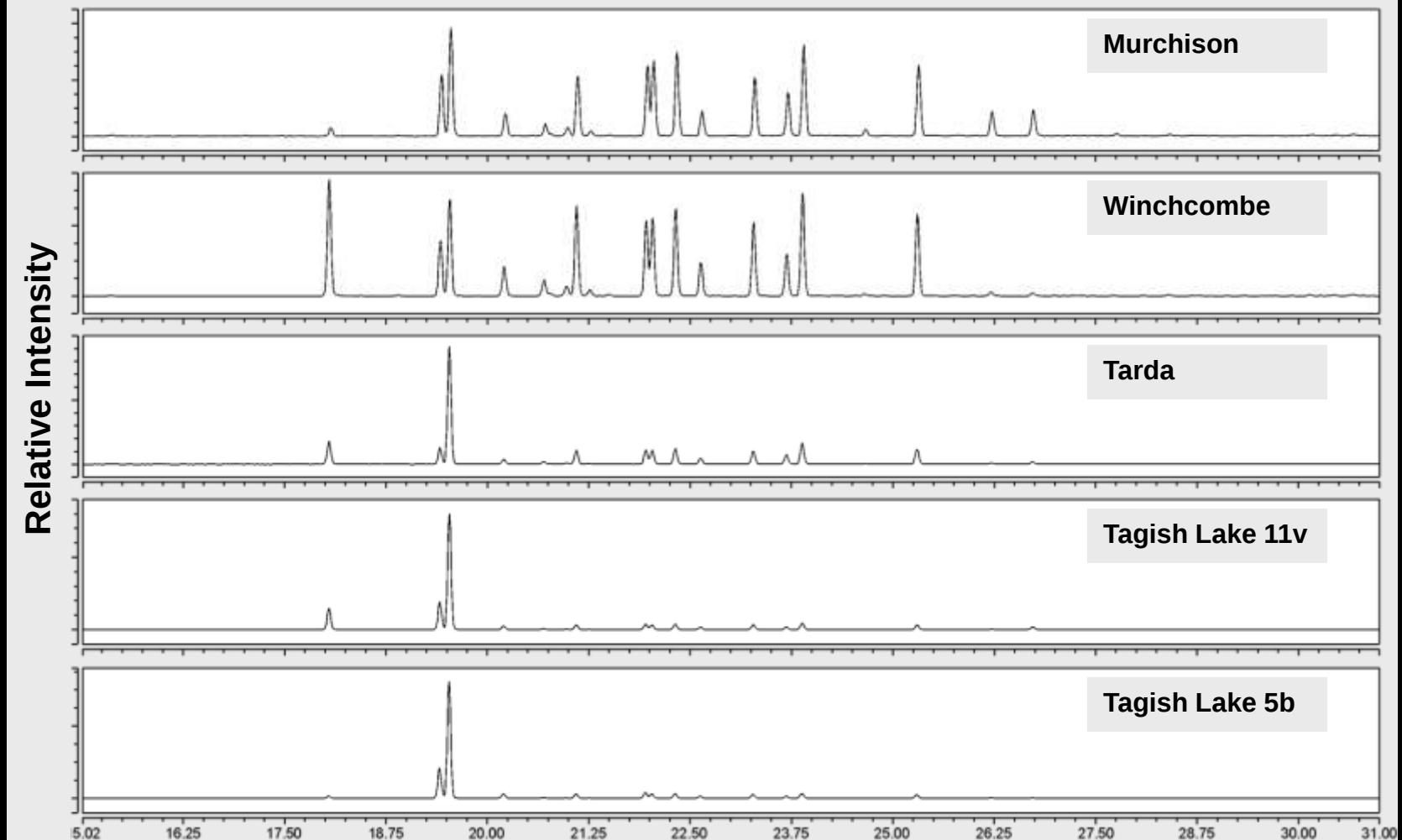
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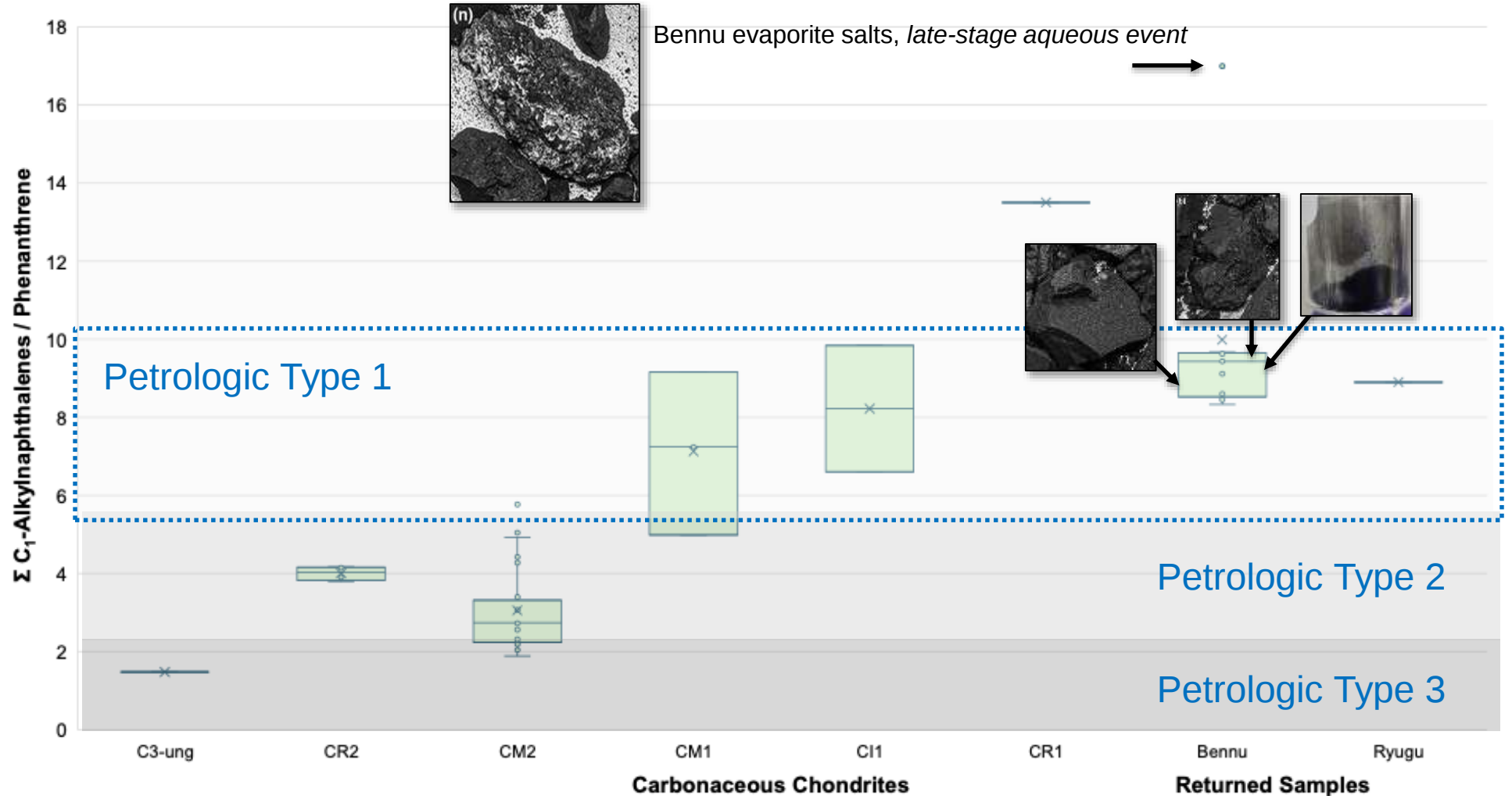
C₄-Alkylbenzene isomer distribution m/z 134.2 $\rightarrow$ 119.1



pyGC-MS: ΣC_1 -MeNp / Ph Ratio

Mojarro et al. (2025) PNAS e2512461122

Increasing aqueous alteration



one-way ANOVA testing of petrologic type 1 (CI1, CM1) type 2 (CR2, CM2), and type 3 (C3ung) indicates there is a significant difference between group means.

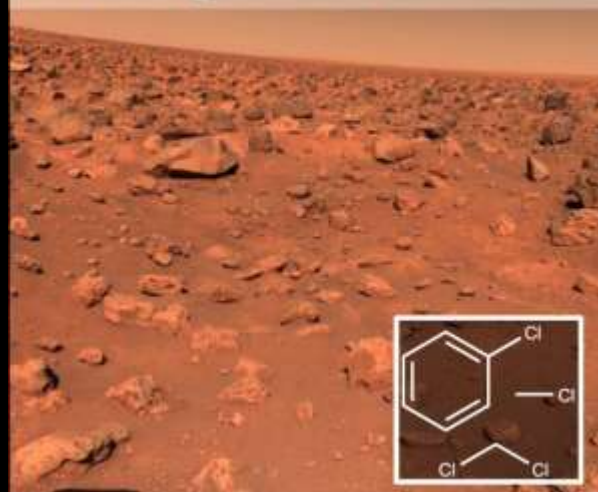
pyrolysis-GC-MS Applied to Martian Organics

Buckner & Mojarro (2026)

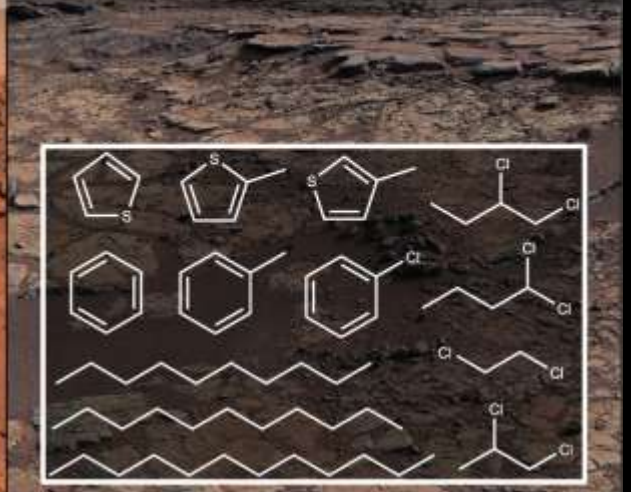


**NASA's Viking Mission May
Have Found Alien Life In 1976
And Accidentally Killed It**

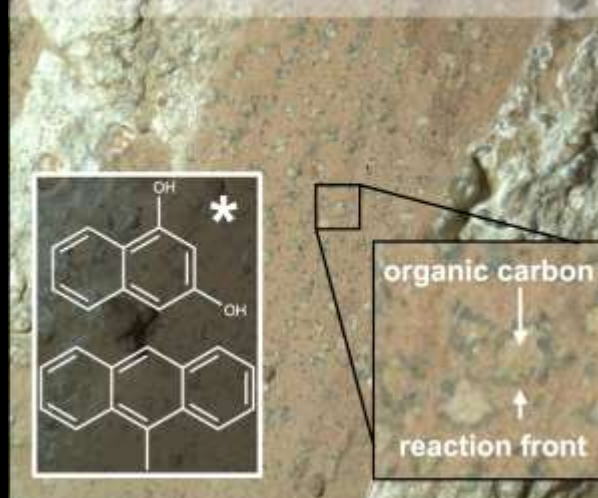
(a) *Utopia Planitia, Mars*
Viking landers



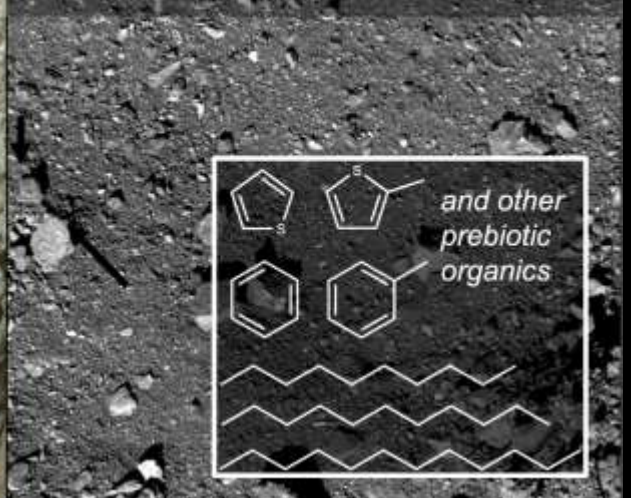
(b) *Gale Crater, Mars*
Curiosity rover



(c) *Jezero Crater, Mars*
Perseverance rover



(d) *Asteroid Bennu*
OSIRIS-REx



Reanalysis of the Viking Results Suggests Perchlorate and Organics at Midlatitudes on Mars

Table 1. Detection of Chlorohydrocarbons in Martian Soils by the Viking TV-GC-MS^a

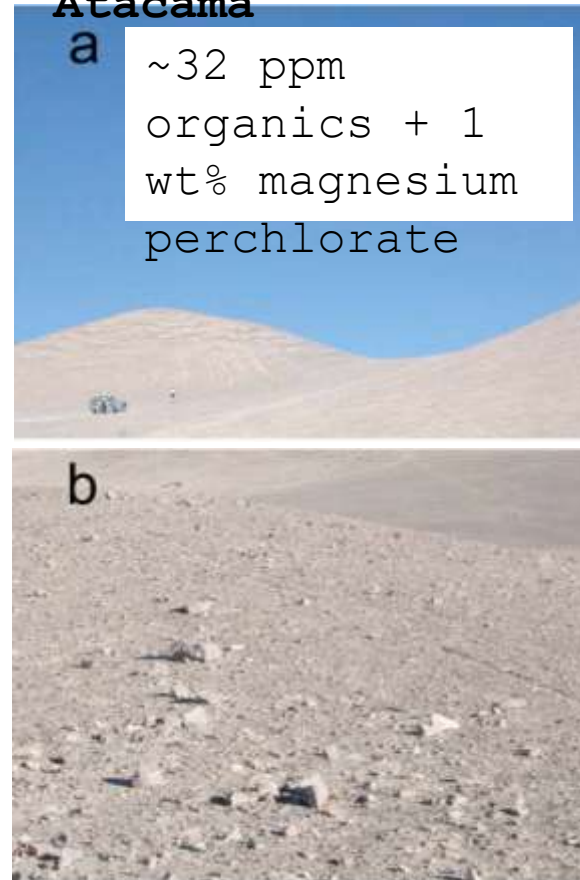
Sample	T (°C)	Injection Mode ^b	Compound	Abundance (ppb)
VL-1				
Blank	500	CO ₂	CH ₃ Cl	ND
Sample 1	200	CO ₂	CH ₃ Cl	15
	500	CO ₂	CH ₃ Cl	ND
Sample 2	200	CO ₂	CH ₃ Cl	ND
	350	CO ₂	CH ₃ Cl	ND
	500	CO ₂	CH ₃ Cl	ND
VL-2				
Blank	500	CO ₂	CH ₂ Cl ₂	ND
Sample 1	200	H ₂	CH ₂ Cl ₂	ND
	350	H ₂	CH ₂ Cl ₂	6–14
	500	H ₂	CH ₂ Cl ₂	6–14
	500	CO ₂	CH ₂ Cl ₂	2–6
Sample 2	50	H ₂	CH ₂ Cl ₂	ND
	200	H ₂	CH ₂ Cl ₂	0.04–0.08
	350	H ₂	CH ₂ Cl ₂	10–20
	500	H ₂	CH ₂ Cl ₂	<4
	500	CO ₂	CH ₂ Cl ₂	20–40

^aFrom *Biemann et al.* [1977]. GC, gas chromatography; MS, mass spectrometry; ND, not determined; TV, thermal volatilization; VL-1 and VL-2, Viking Landers 1 and 2, respectively.

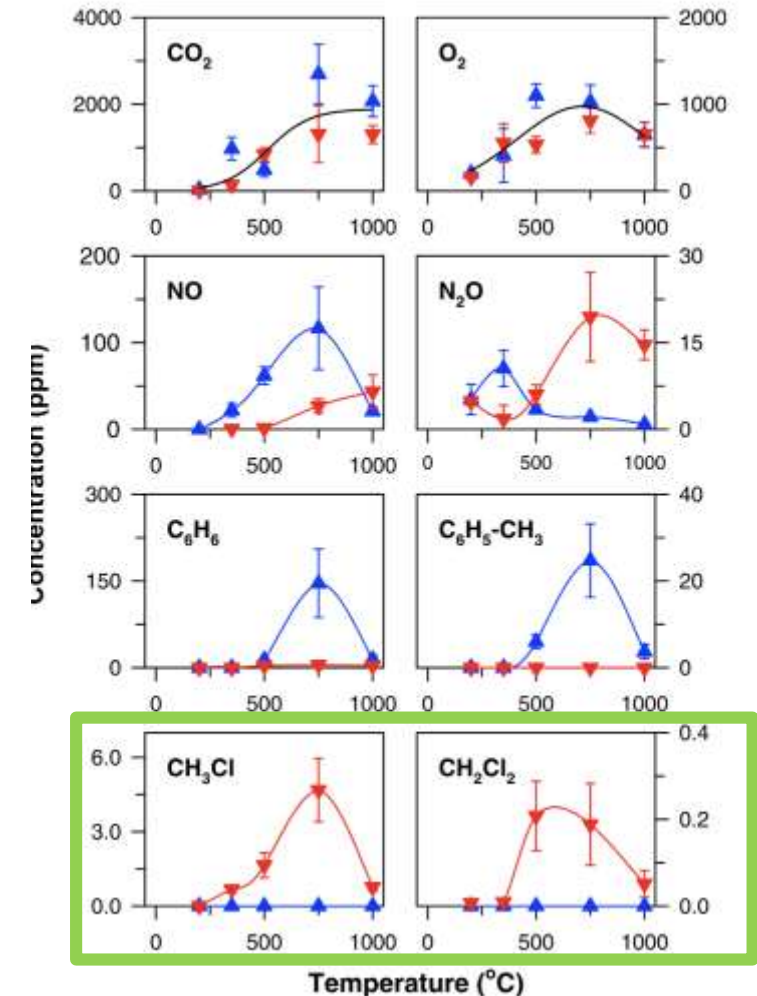
^bGases and volatiles released by TV are swept away from the oven into the gas chromatograph by a stream of either ¹³CO₂ or H₂.

Yungay region

Atacama

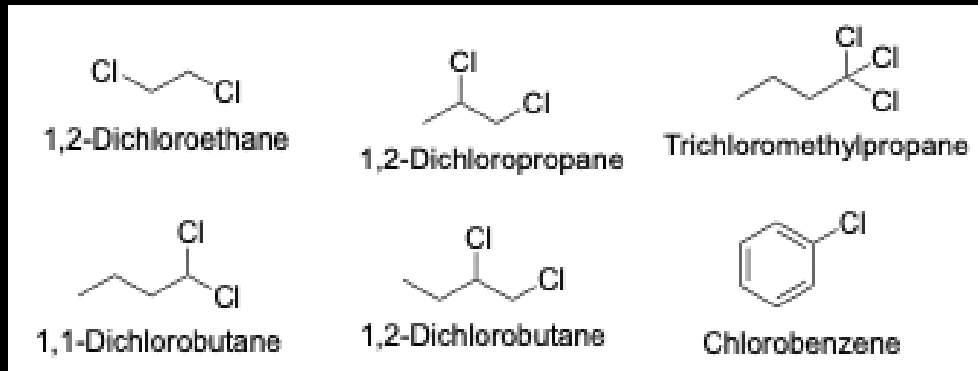


from Navarro-Gonzalez et al., 2010



SAM Confirmed Chlorinated Compounds Are Indigenous

- The Sample Analysis at Mars (SAM) instrument aboard the Curiosity rover is a more sophisticated *pyGC-MS*.
- Chlorinated compounds detected above background levels:



Freissinet et al. 2015; Szopa et al. 2020

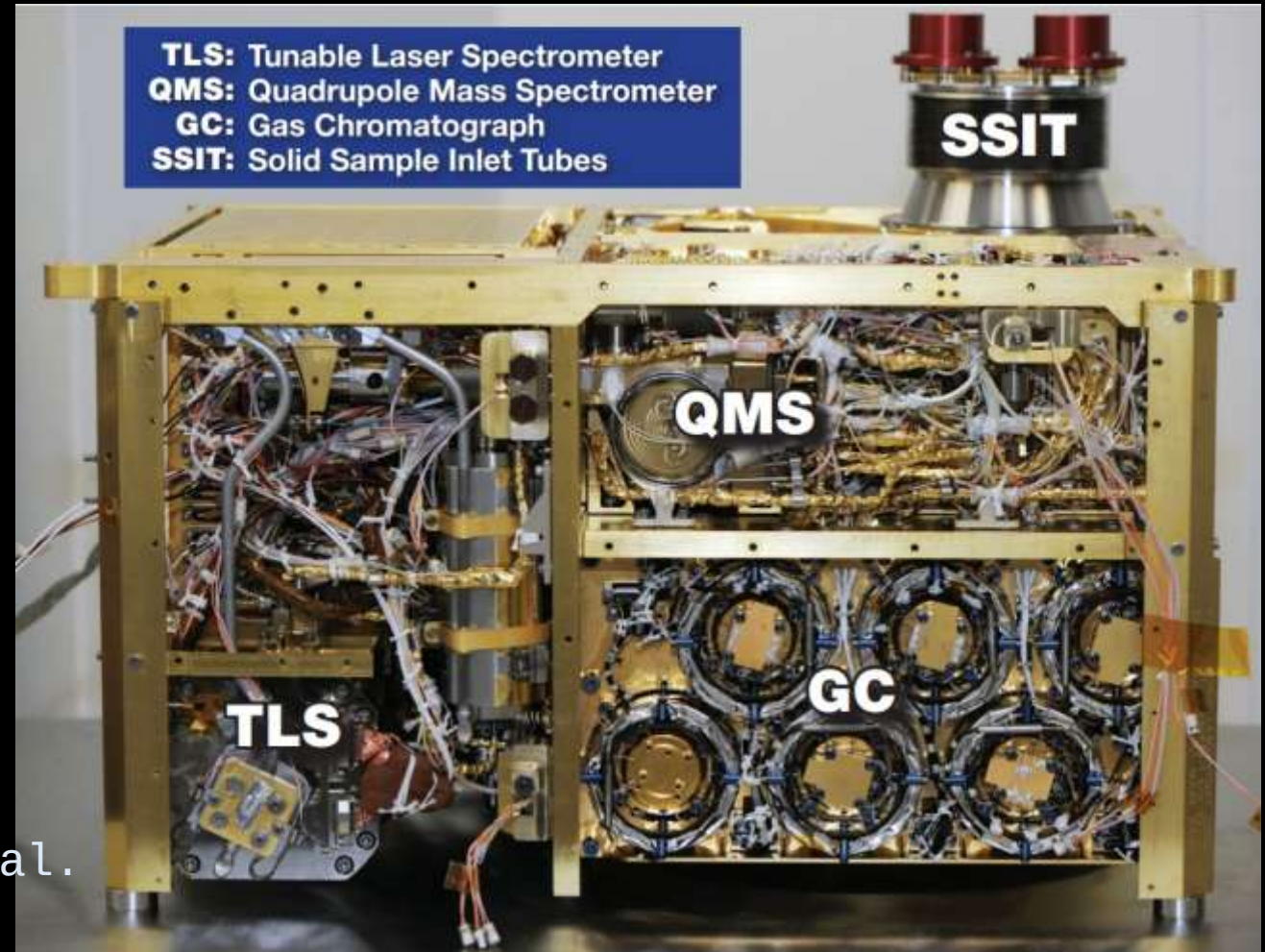
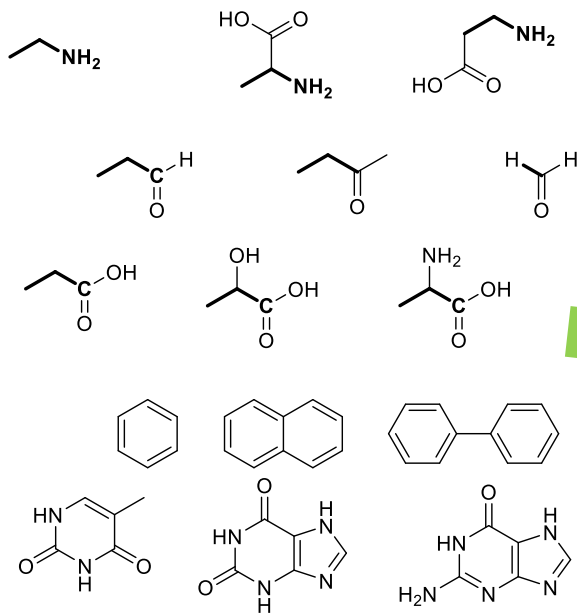


Image credit: NASA

Origin of the organic compounds (abiogenic or biogenic) remains uncertain. Determining the source of organic matter on Mars may ultimately require sample return to Earth laboratories

Organic Compounds in Martian Regolith

Soluble Organic Matter (SOM) / Bitumen



Solvent-extractable labile fraction more susceptible to degradation.

Sample

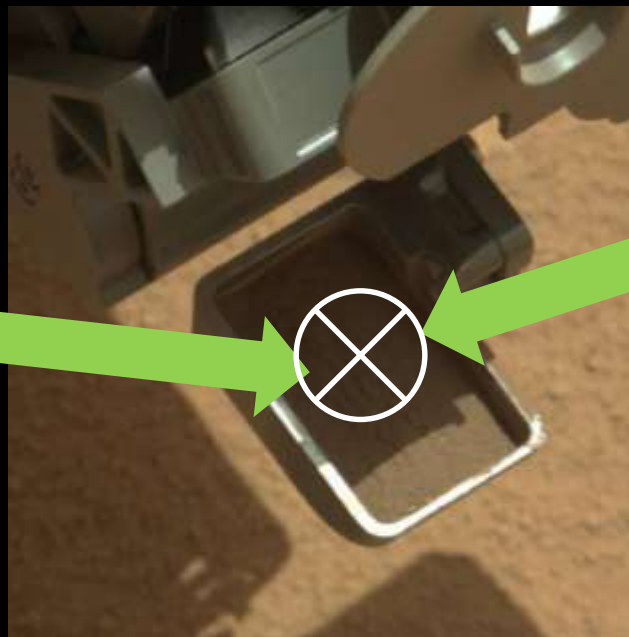
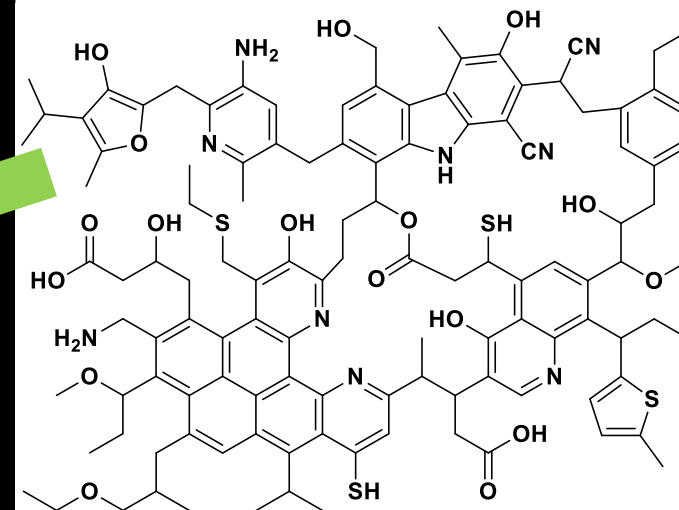


Image credit: NASA JPL

Insoluble Organic Matter (IOM) / Kerogen



Refractory fraction can persist through diagenesis on geological timescales.

(IOM from Glavin et al., 2013)

The Martian Meteorite Record

- Shergottites (basalts), Nakhrites (clinopyroxenite), Chassignites (dunite) plus ALH 84001 (orthopyroxenite) and NWA 7034 (breccia) - SNC meteorites.
- Martian origin of SNC meteorites determined by comparison of trapped gas composition to Viking atmospheric measurements and oxygen isotopes.
- Delivered by impacts without geological context and terrestrial residence time increases contamination risk.
- Most SNCs are relatively young (< 2.4 Ga) igneous rocks from volcanic activity.



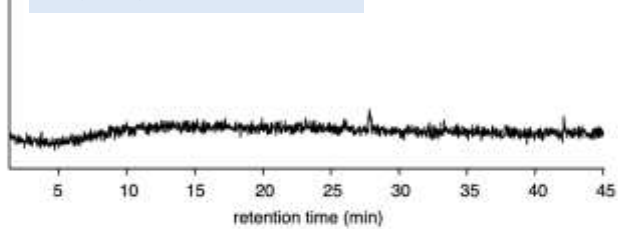
Image credit: NASA JSC

Electron microscopy revealed chain structures resembling living organisms in meteorite fragment ALH84001.

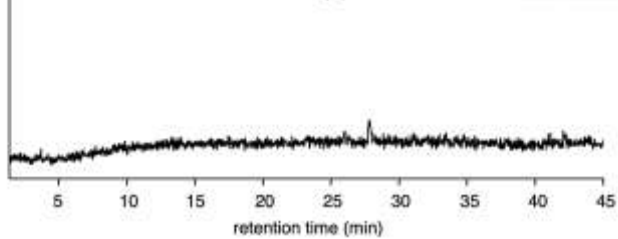
Organic Compounds in SNCs

Pyrolysis @ 610 °C

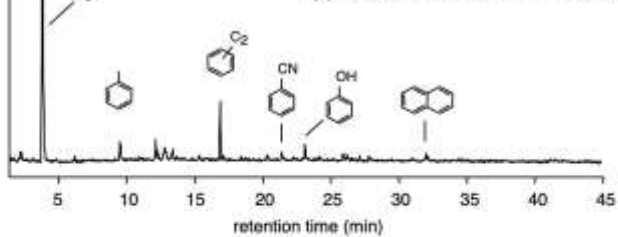
(a) procedural blank



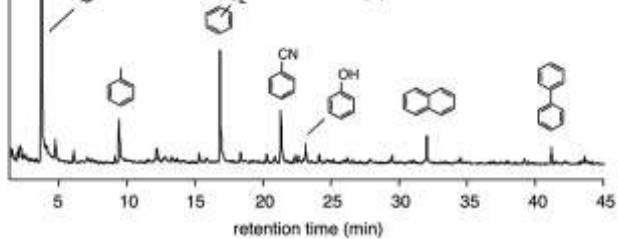
(b) solvent extracted ALH 84001



(c) solvent extracted EET A79001



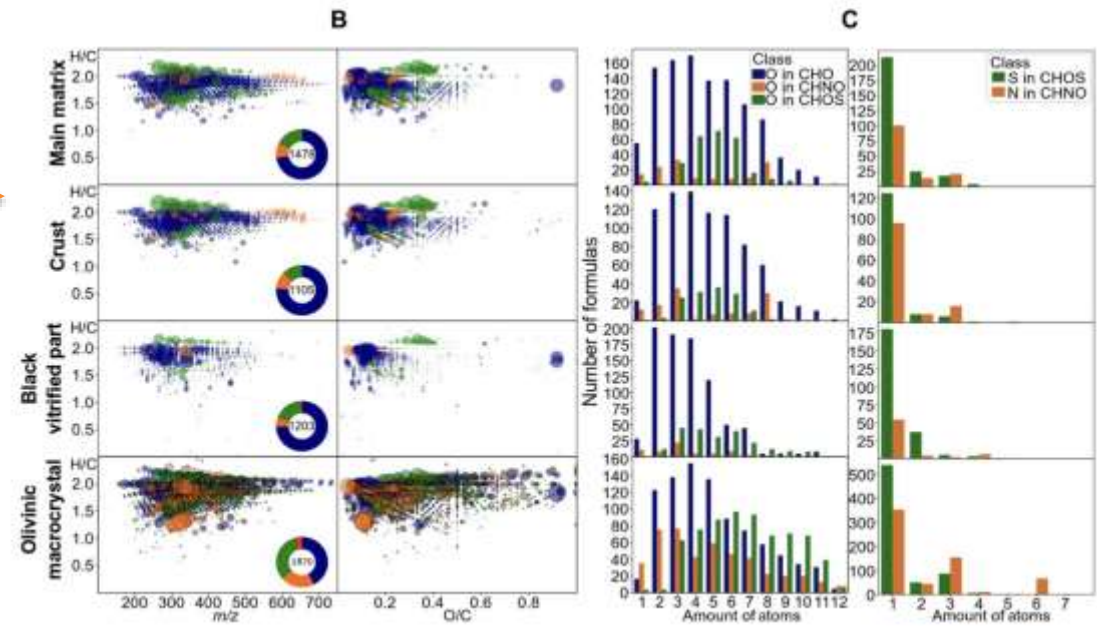
(d) solvent extracted Nakhla



High molecular weight organic matter in Martian meteorites revealed by pyGC-MS

Complex soluble organic matter in Tissint Martian meteorite samples including carboxylic acids from $C_3 - C_7$

Sephton et al., 2002



Schmitt-Kopplin et al., 2023

SNC Coordinated Analysis

Melt&freeze dry ~1 L of blue ice (valuable control for Antarctic meteorites)

Tarda sand (useful control for Tissint)

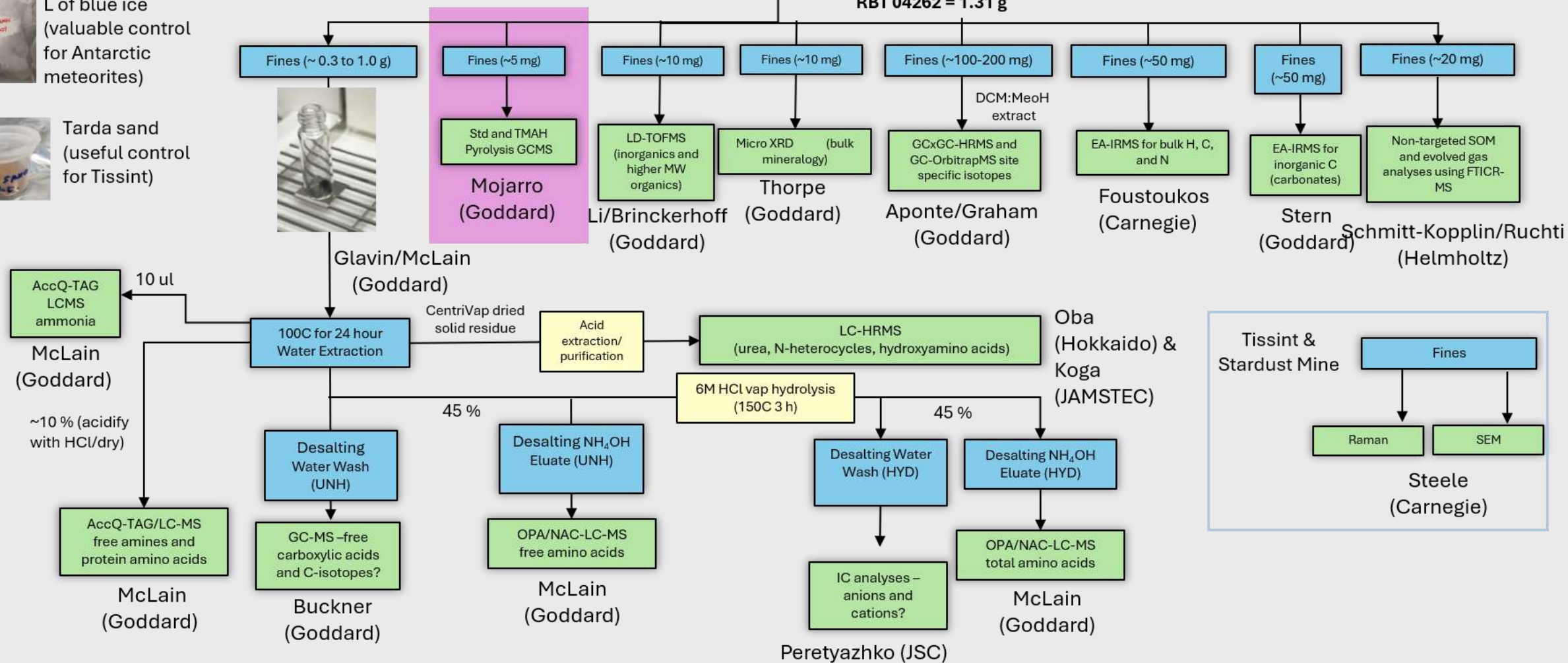


Crushing and photo documentation. ALH 840001 and RBT 04262 powdered using quartz mortar and pestle inside HEPA filtered flow bench (ISO 5) in white room at Goddard by Angel Mojarro on 5/5/25.

Post-crush sample mass transferred to glass vial:

ALH 84001 = 1.19 g

RBT 04262 = 1.31 g



Stepwise Pyrolysis of SNC Meteorites

Characterize different carbon reservoirs
SOM/bitumen and IOM/Kerogen.

I. 50 – 300 °C: Thermal desorption

- I. Free and adsorbed compounds
- II. Terrestrial contamination

II. 600 °C: Pyrolysis

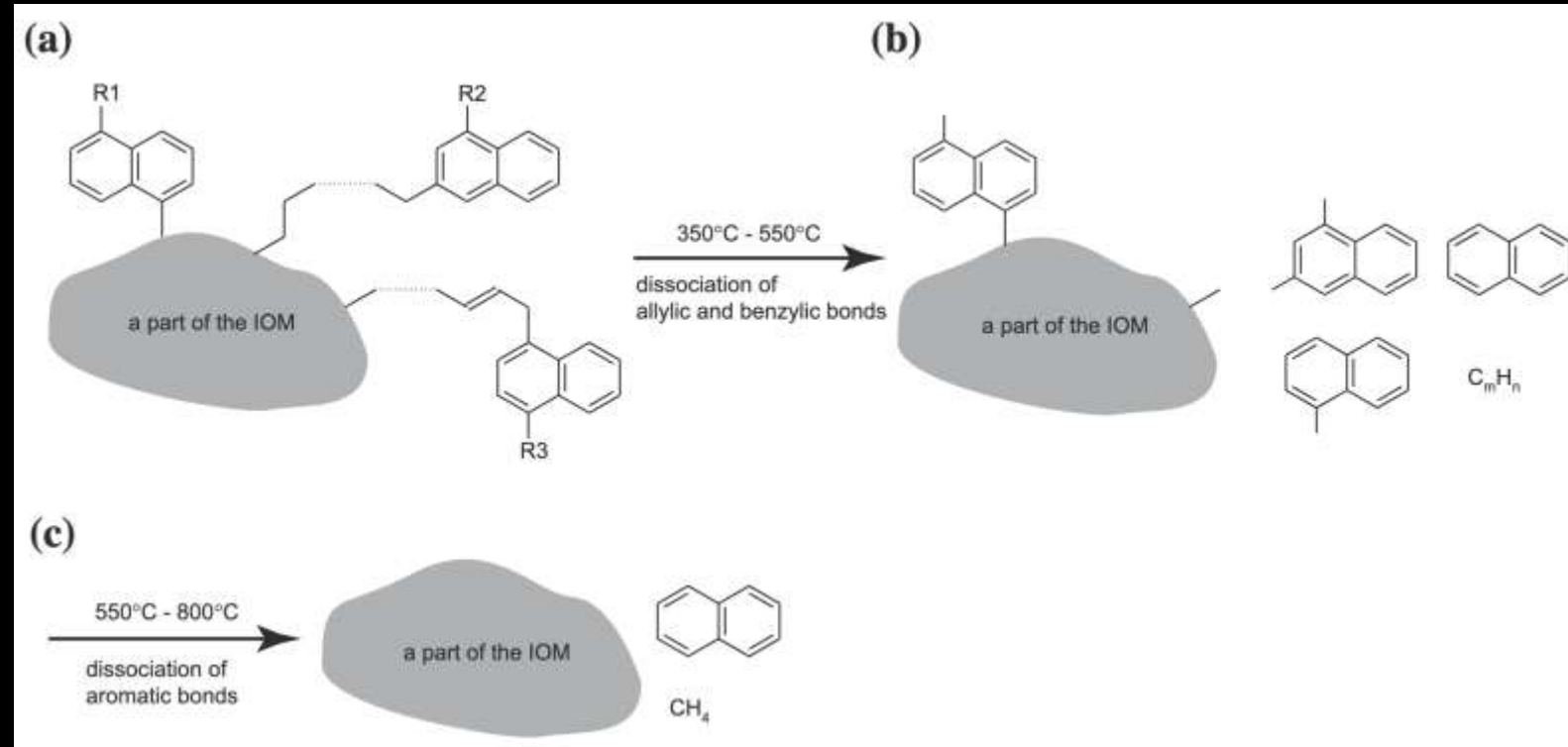
- I. Macromolecular organic matter
- II. Terrestrial contamination

III. 800 °C

- I. Refractory organics

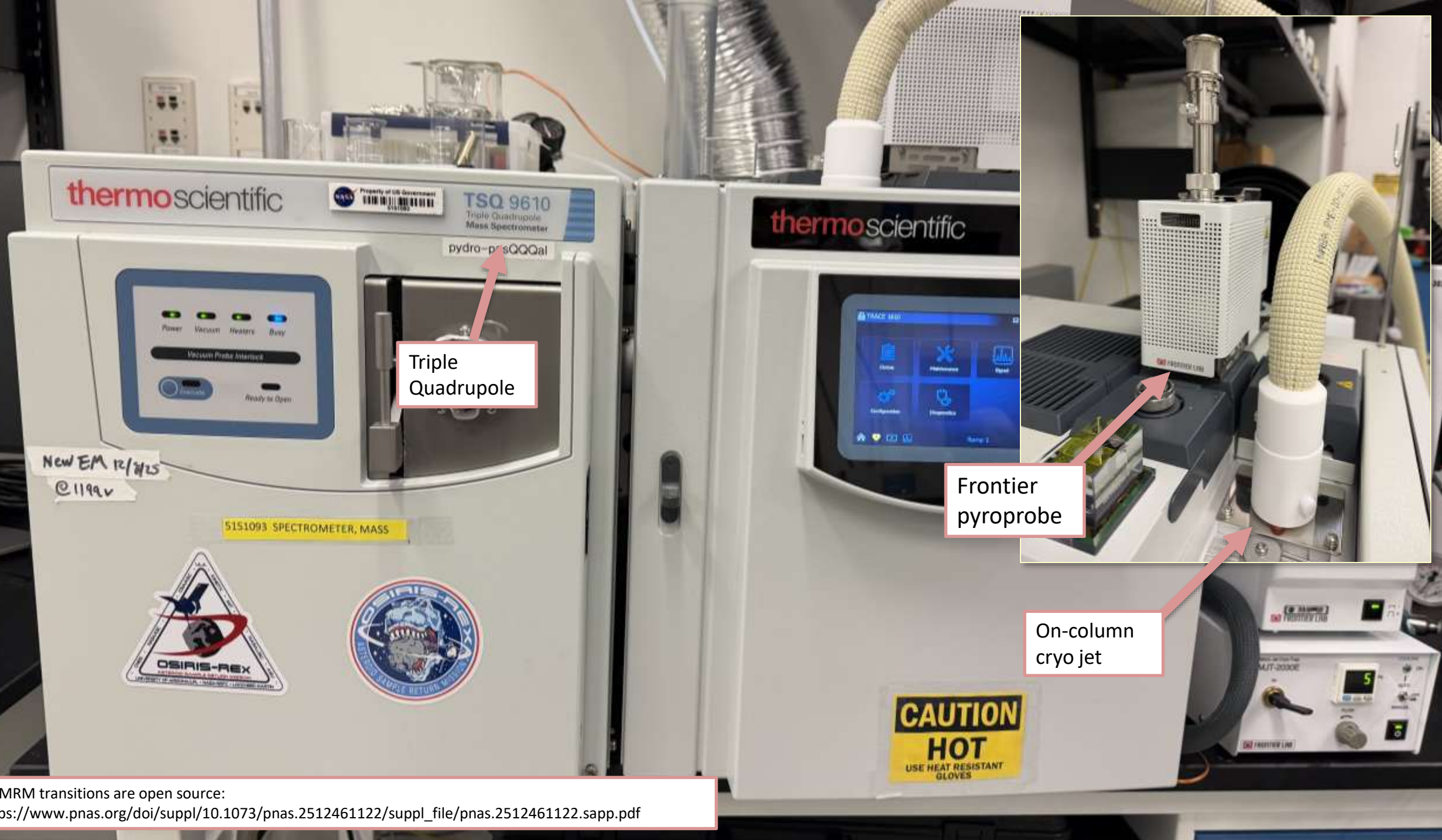
IV. 1000 °C

- I. Graphitic carbon



Gradual and stepwise pyrolysis of insoluble organic matter from the Murchison meteorite reveals structural and isotopic heterogeneity.

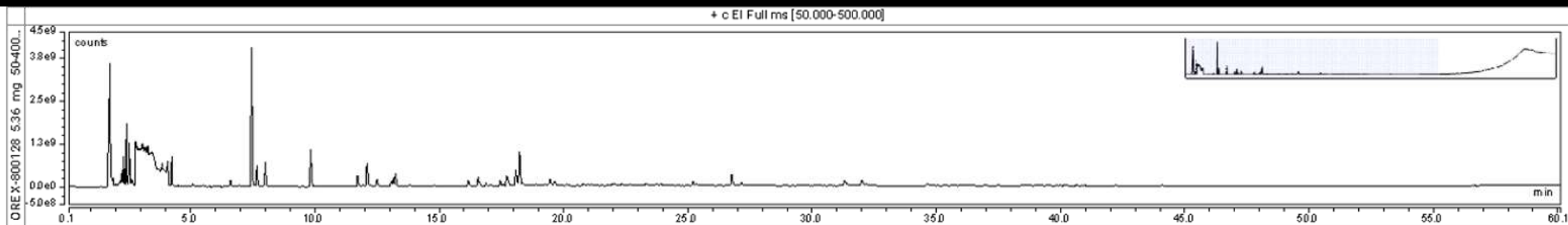
Okumura & Mimura 2011



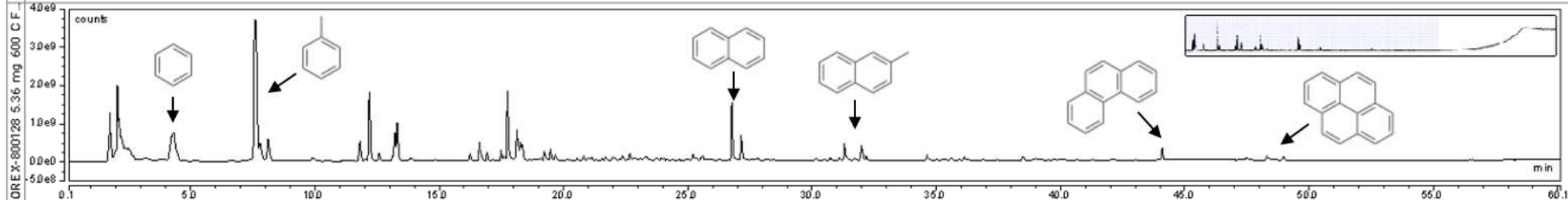
All MRM transitions are open source:
https://www.pnas.org/doi/suppl/10.1073/pnas.2512461122/suppl_file/pnas.2512461122.sapp.pdf

SNC: Tissint (pristine)

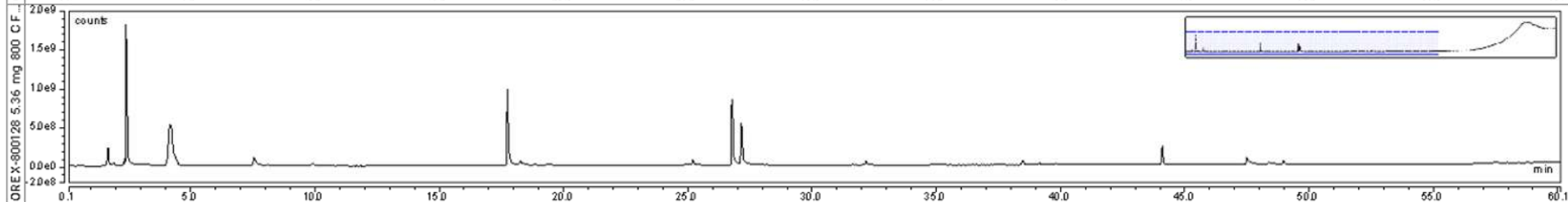
50 – 300 °C



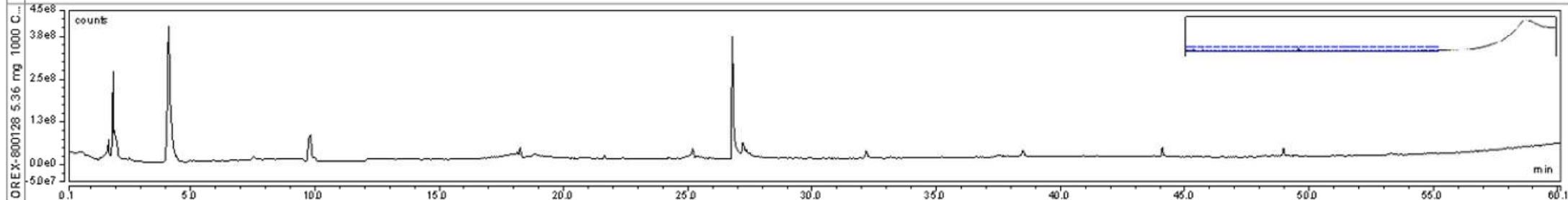
600 °C



800 °C



1000 °C



free volatile macromolecular

Benzene (m/z 78.1 $\rightarrow$ 77.1)

50 – 300 °C

600 °C

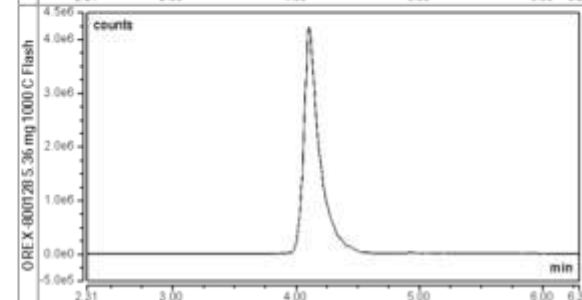
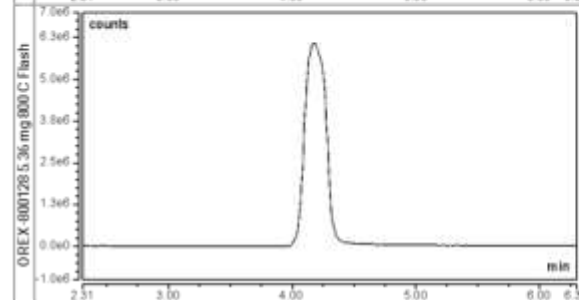
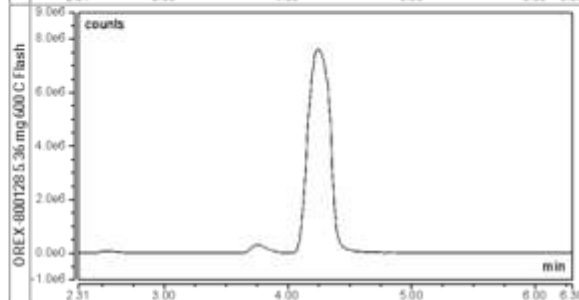
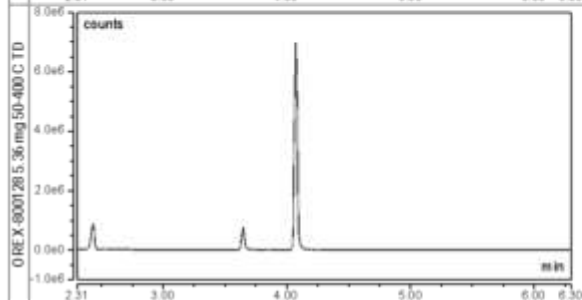
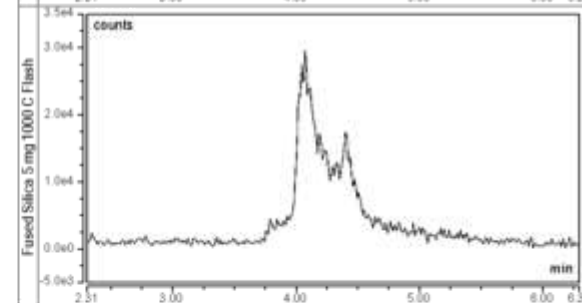
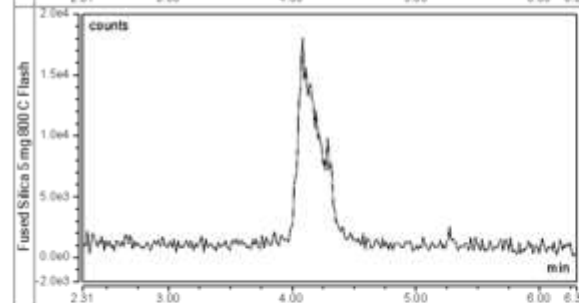
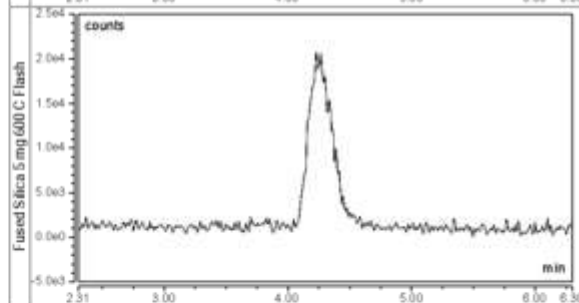
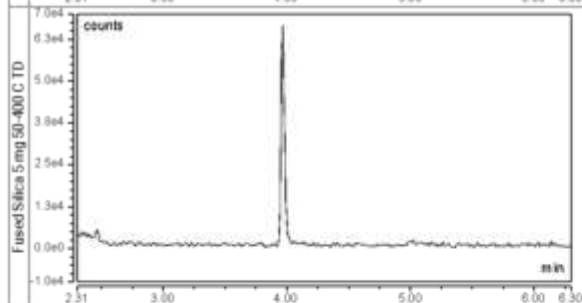
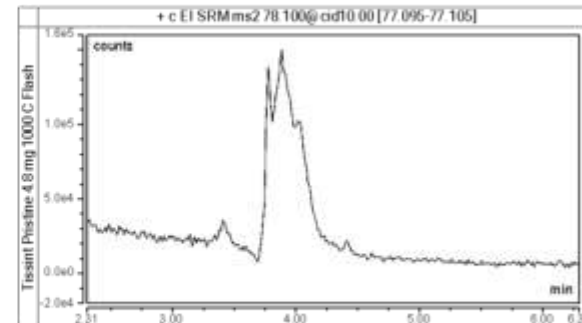
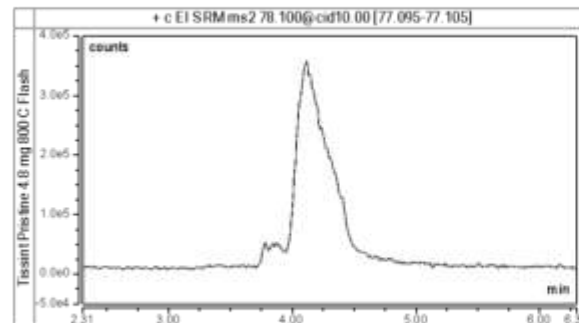
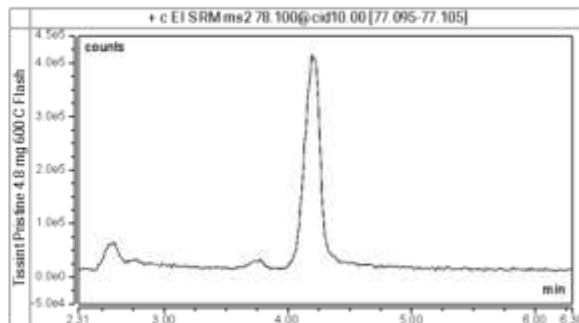
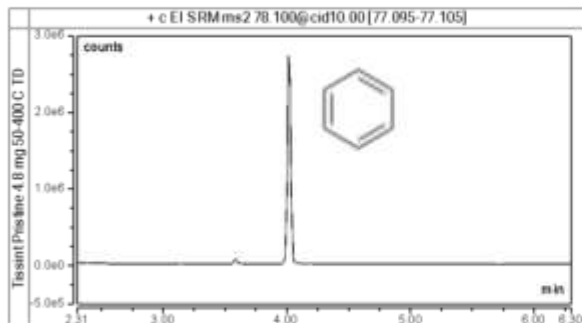
800 °C

1000 °C

Tissint

Blank

OREX



Thiophene (m/z 84.1 → 58)

50 – 300 °C

600 °C

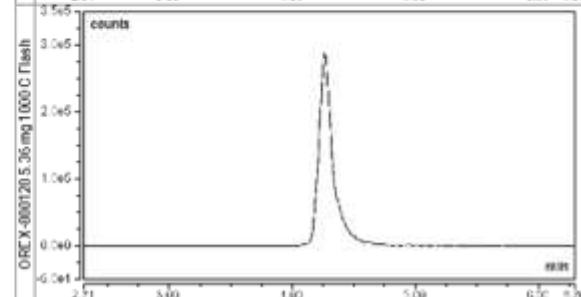
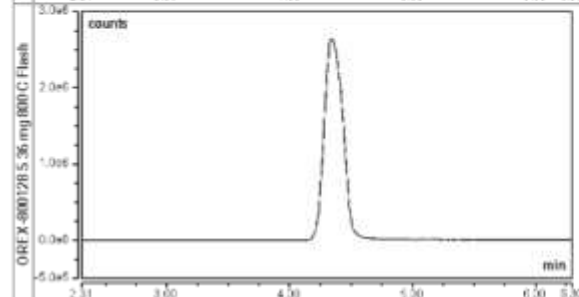
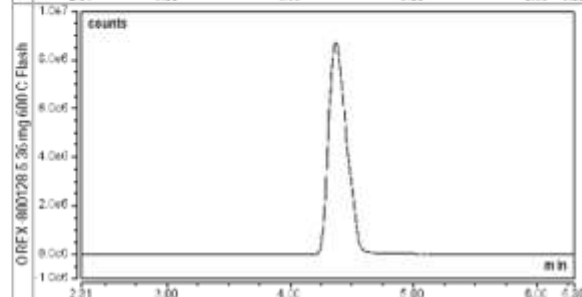
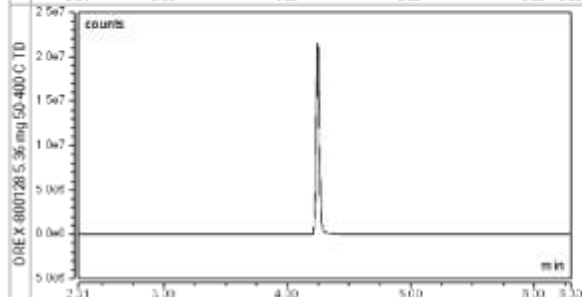
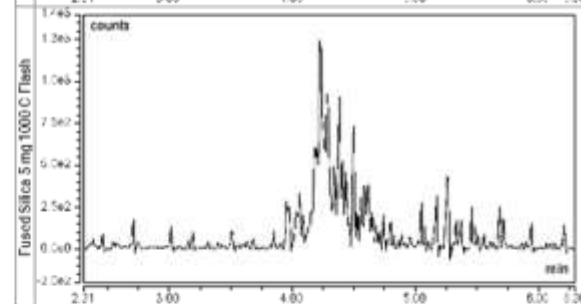
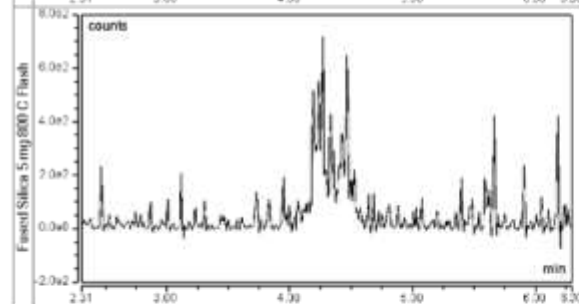
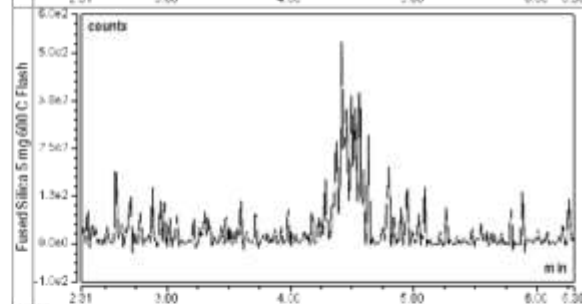
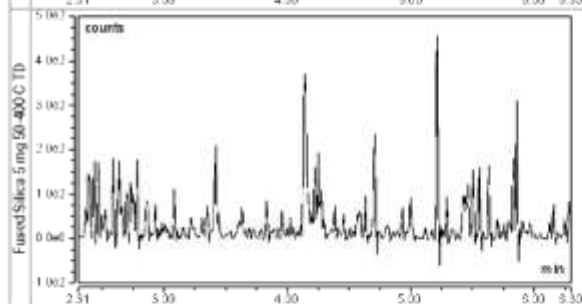
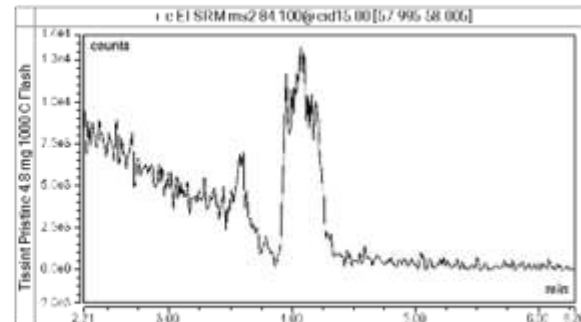
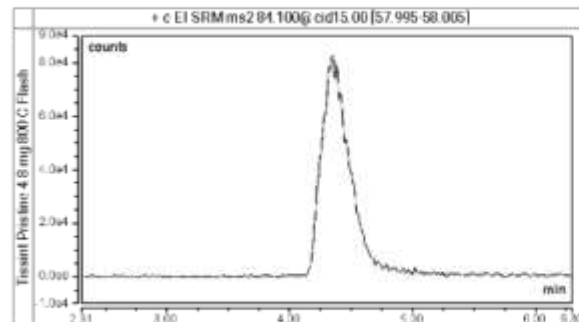
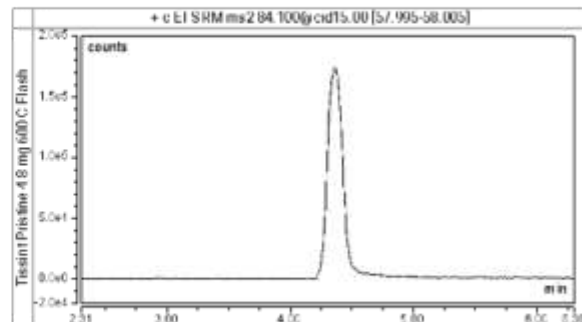
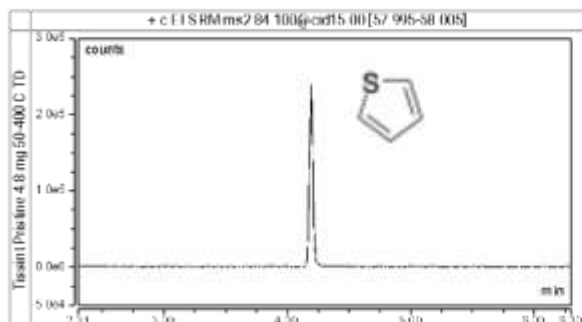
800 °C

1000 °C

Tissint

Blank

OREX



Naphthalene (m/z 128.1 $\rightarrow$ 127.1)

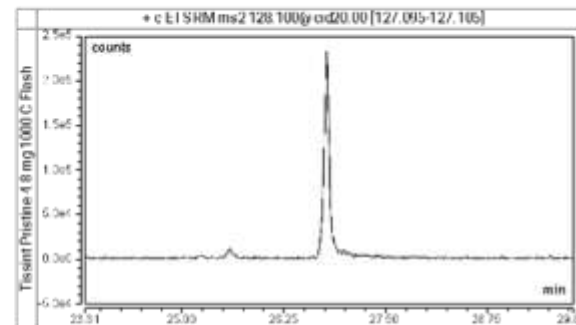
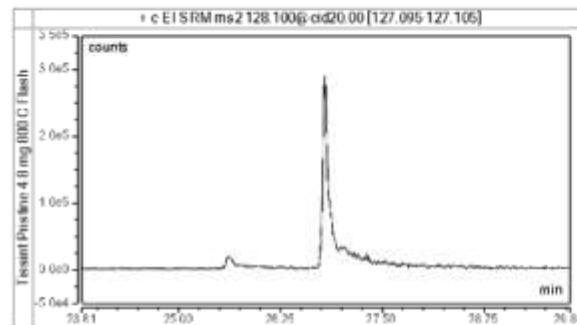
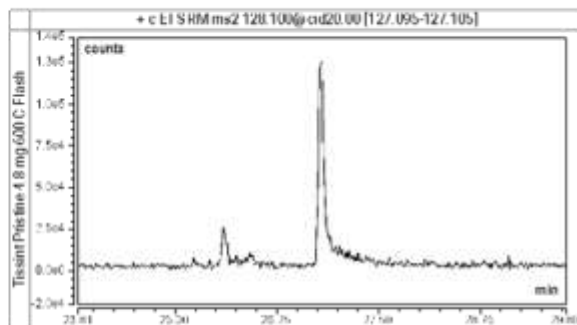
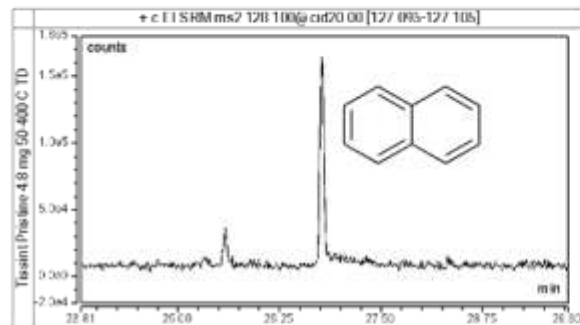
50 – 300 °C

600 °C

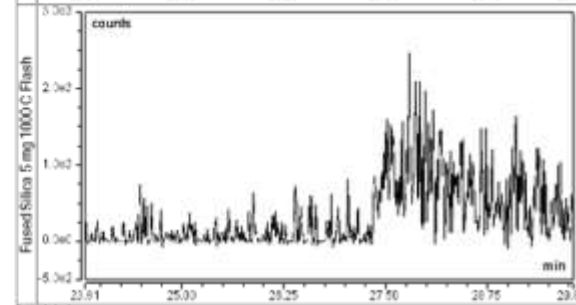
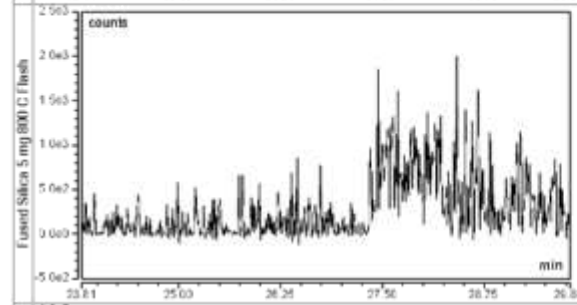
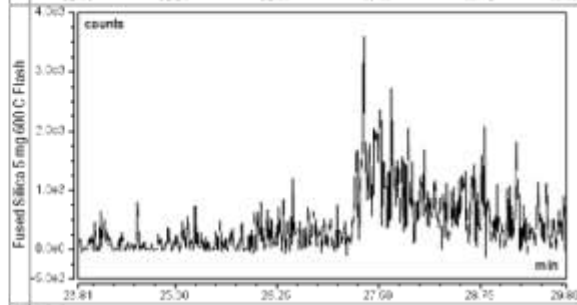
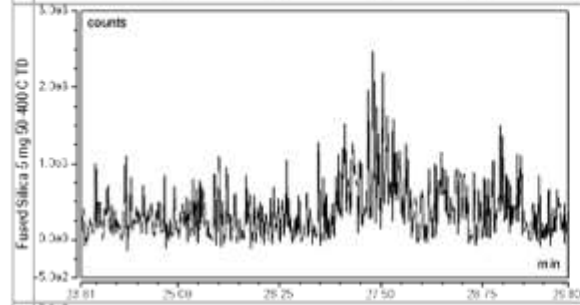
800 °C

1000 °C

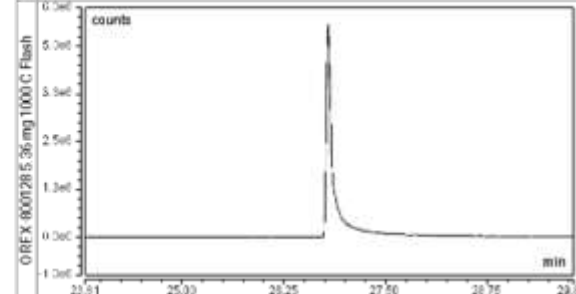
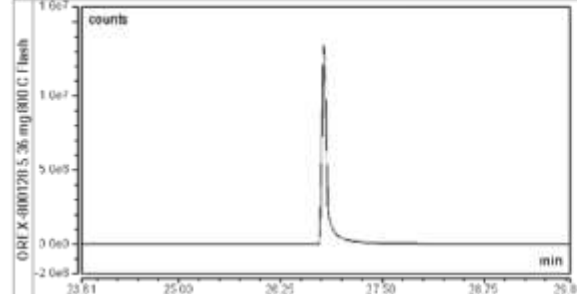
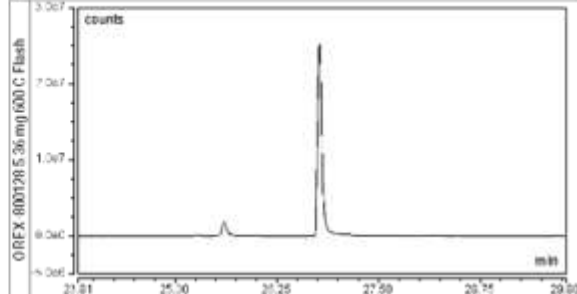
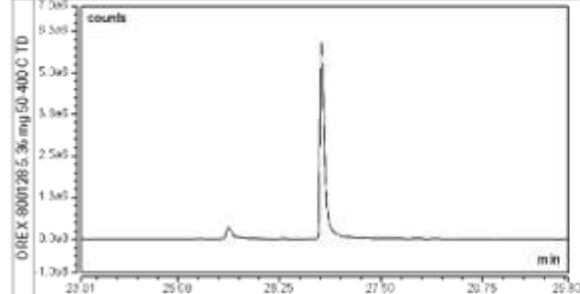
Tissint



Blank

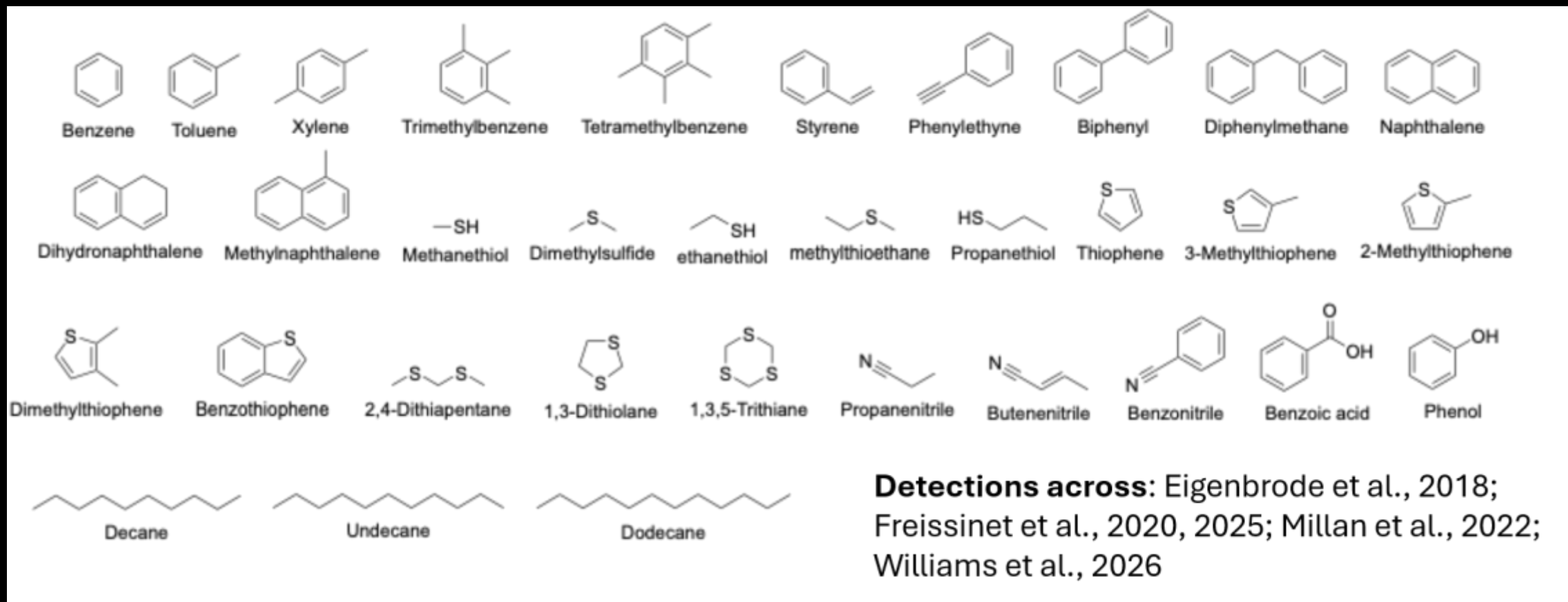


OREX



Work in Progress

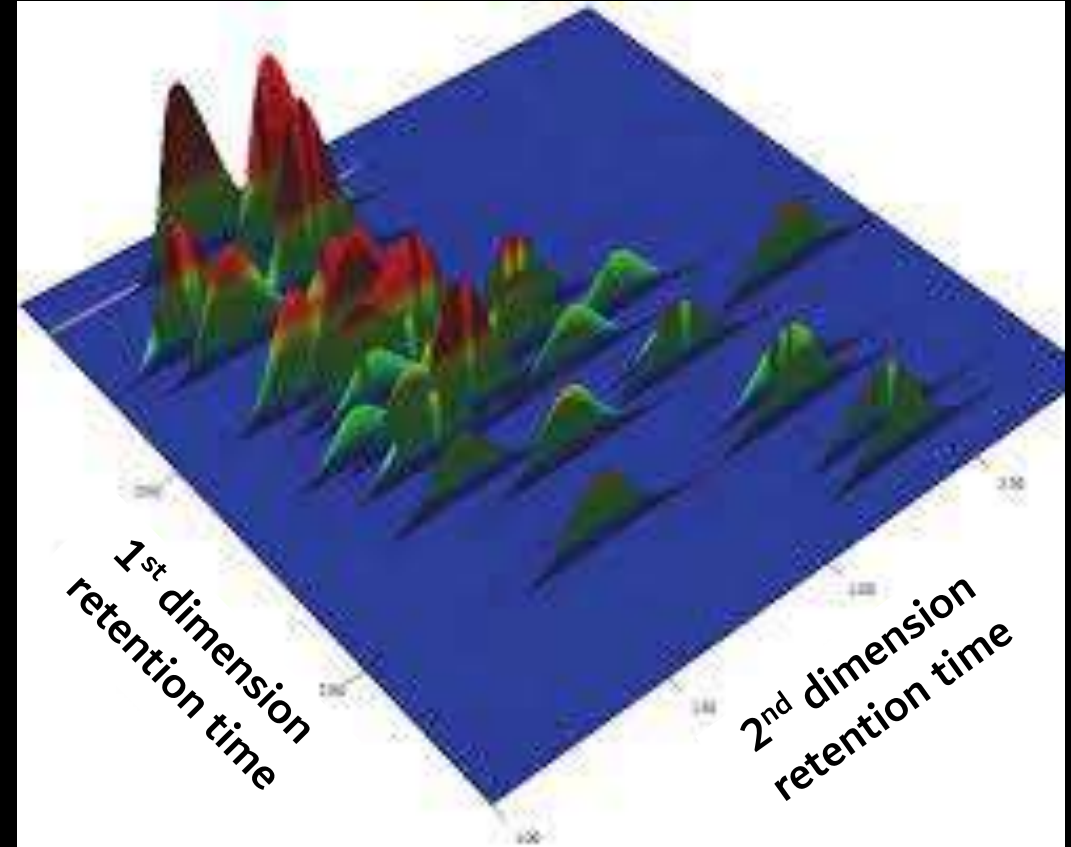
Organic matter preserved in 3-billion-year-old mudstones at Gale crater, Mars



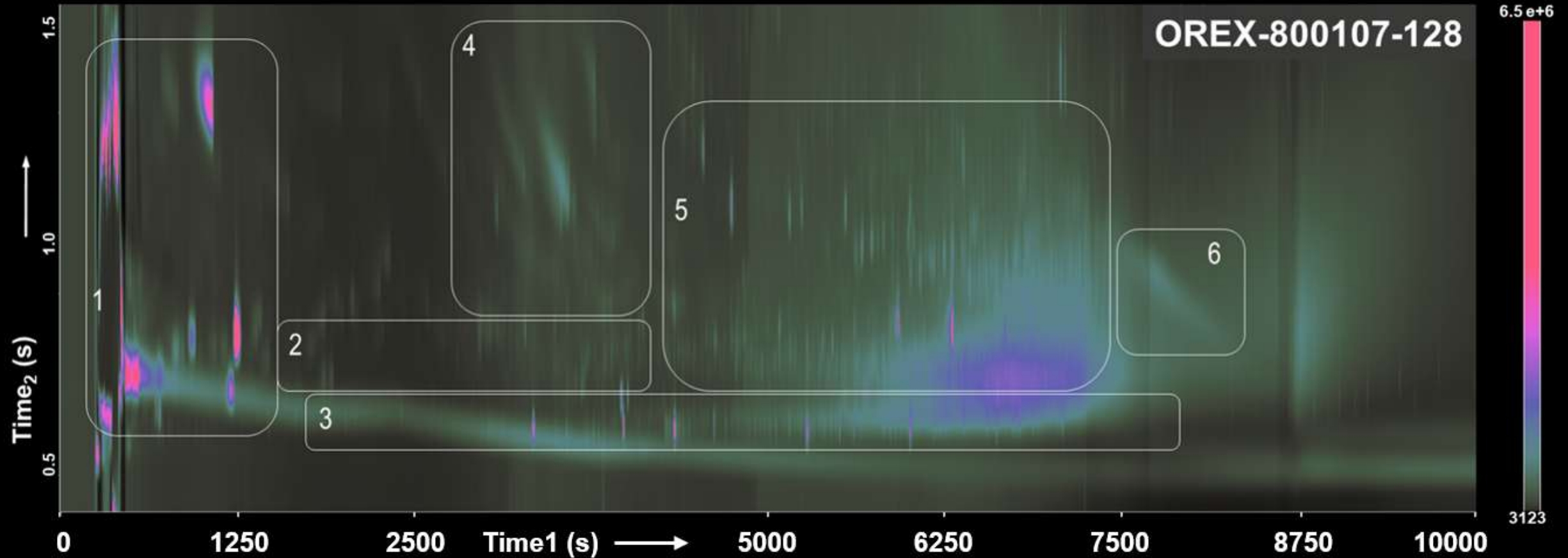
Soluble organics: GC×GC-HRMS Analyses

Two-dimensional gas chromatography allows for an untargeted search of soluble organic species belonging to compound classes of varying polarities.

Solvent extracts were analyzed by an EI-TOF-HR-MS (25,000 $m/\Delta m$ at 219 Da) to determine the accurate mass molecular formula.

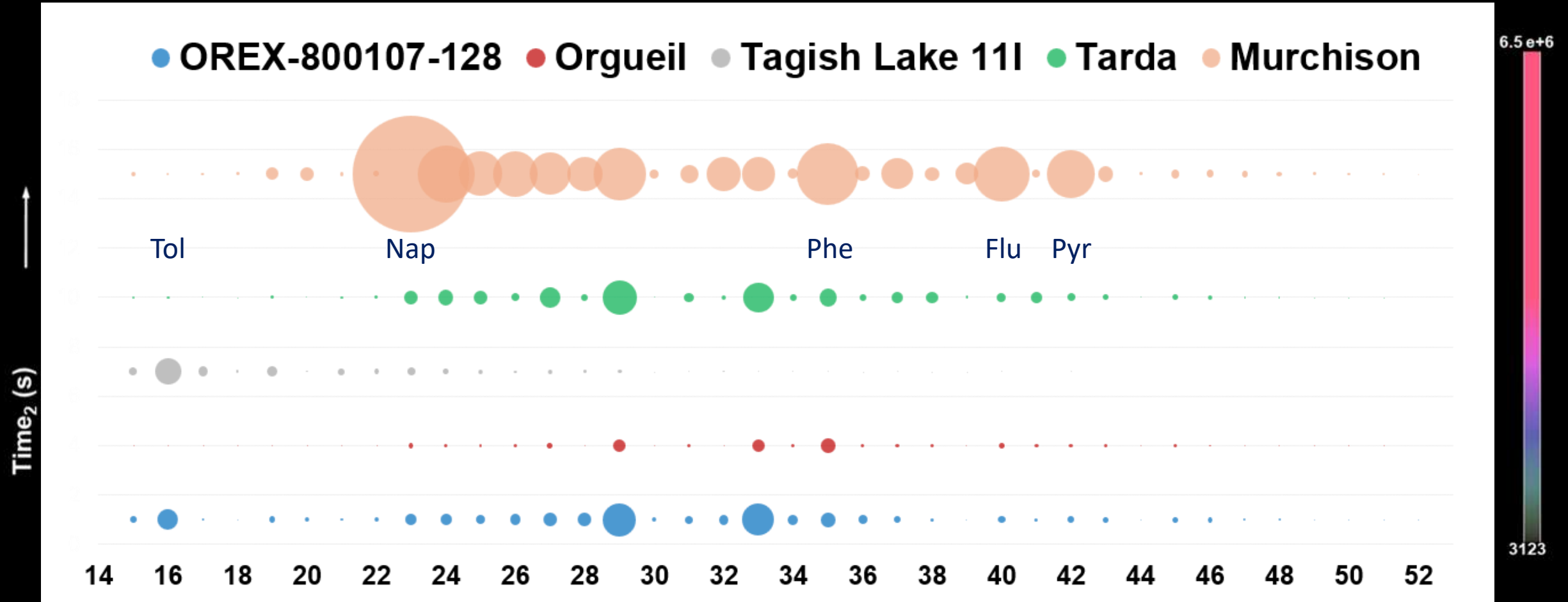


Soluble Organics in Bennu



We found 38 polycyclic aromatic hydrocarbons, and many other N-, O-, and S-containing aromatic species.

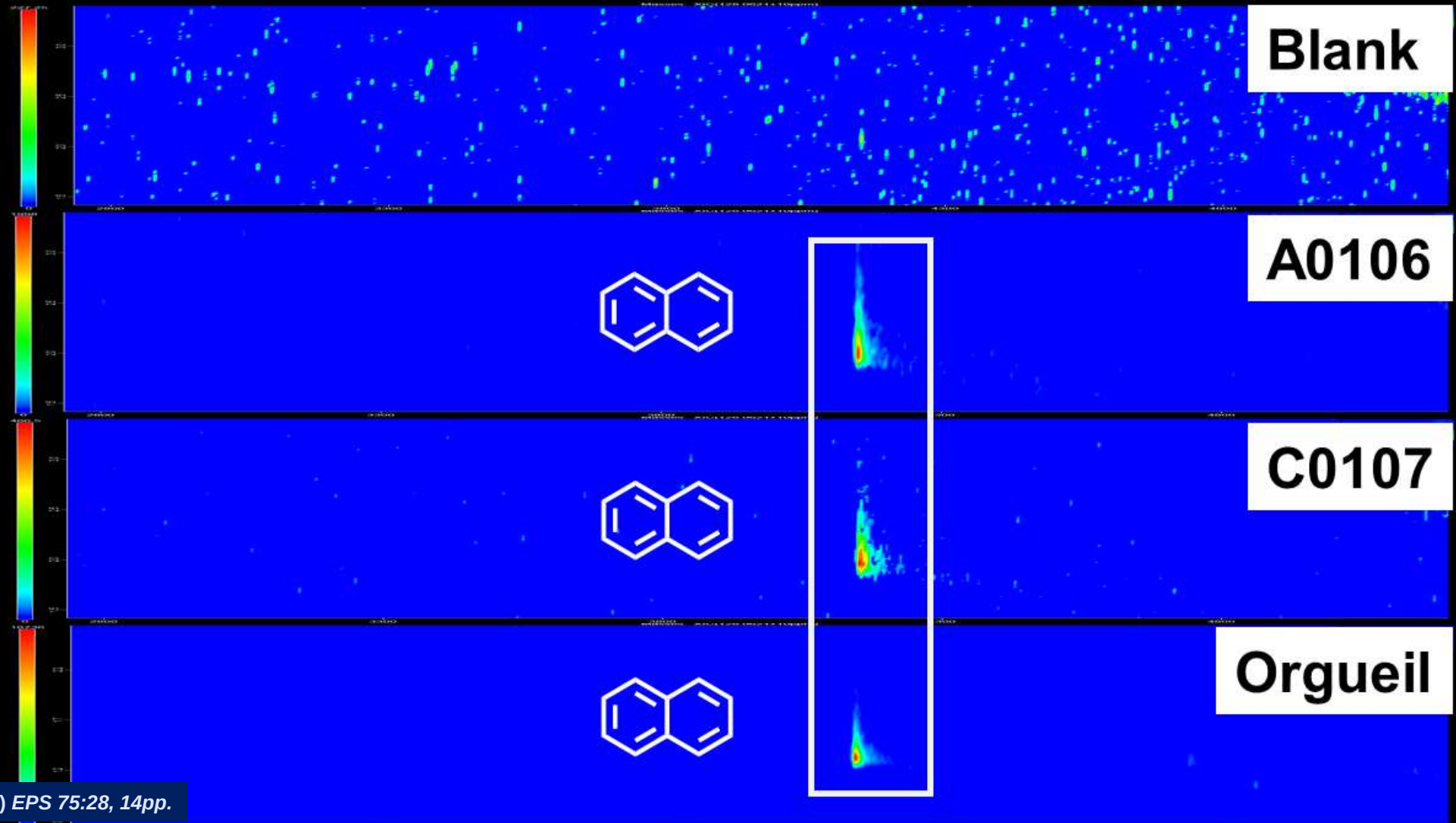
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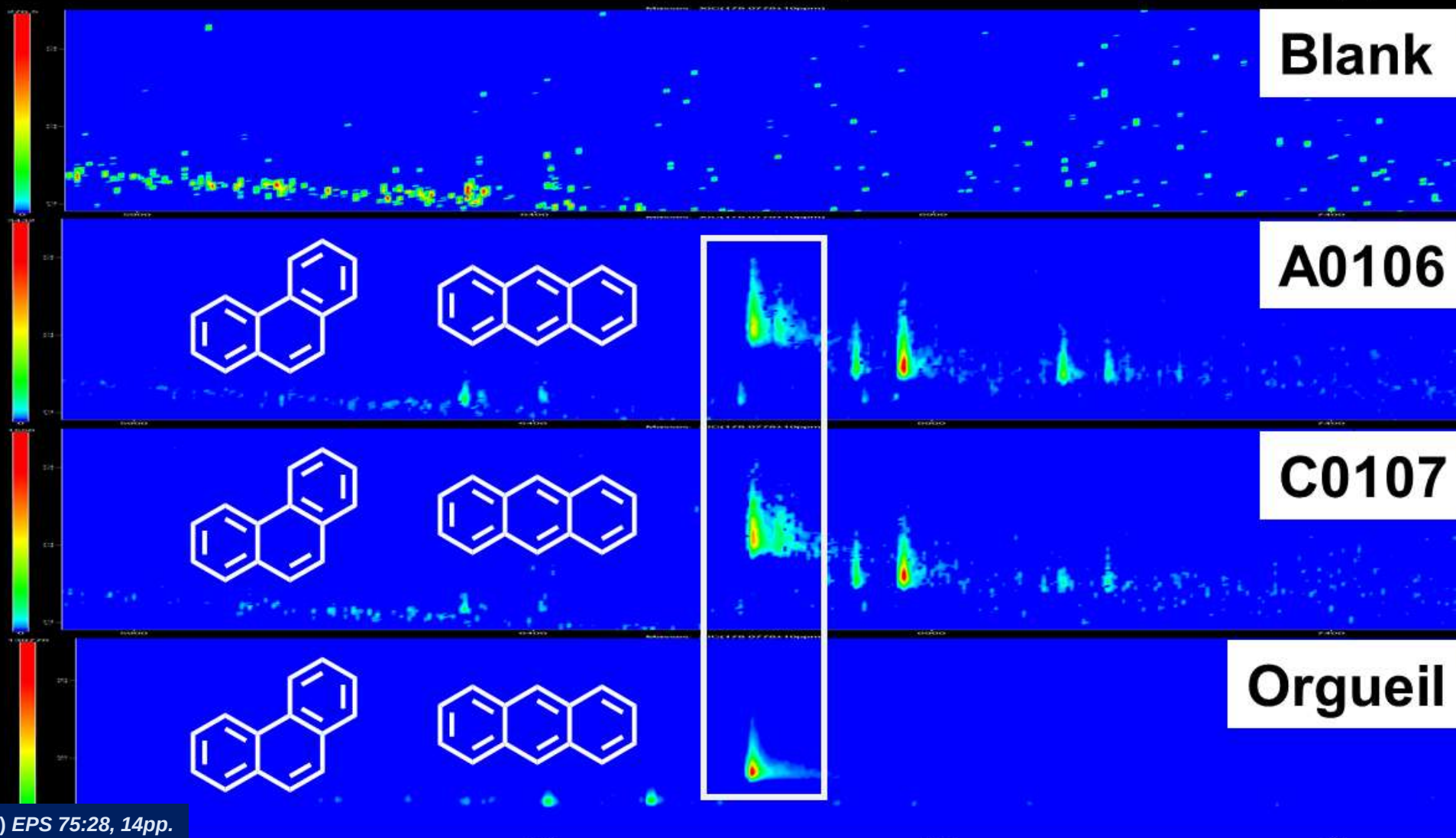
Ryugu: Naphthalene

Naphthalene, m/z 128.0621 $\pm$ 10 ppm



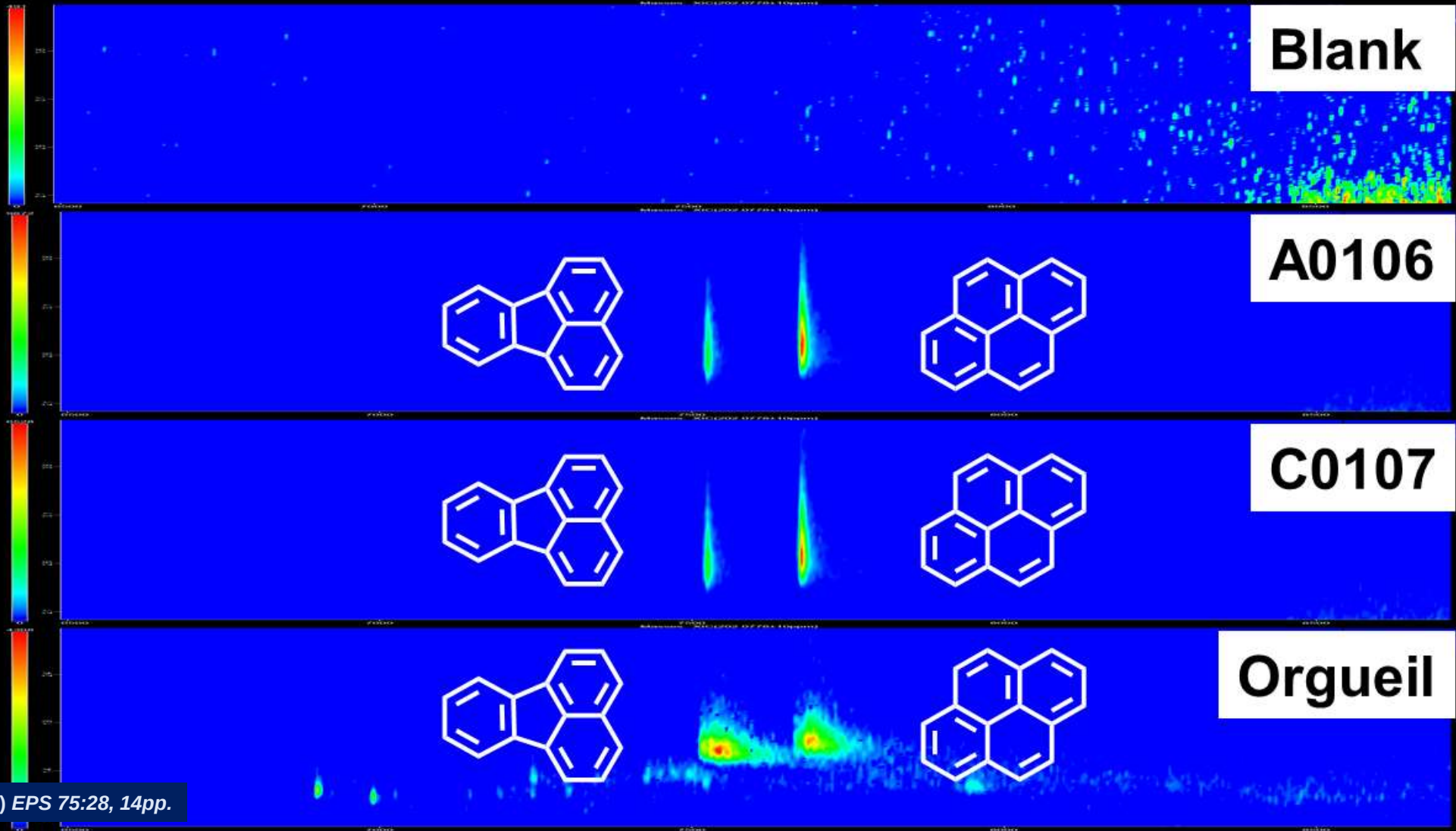
Ryugu: Phenanthrene and Anthracene

Phenanthrene and anthracene, m/z 178.0778 $\pm$ 10 ppm



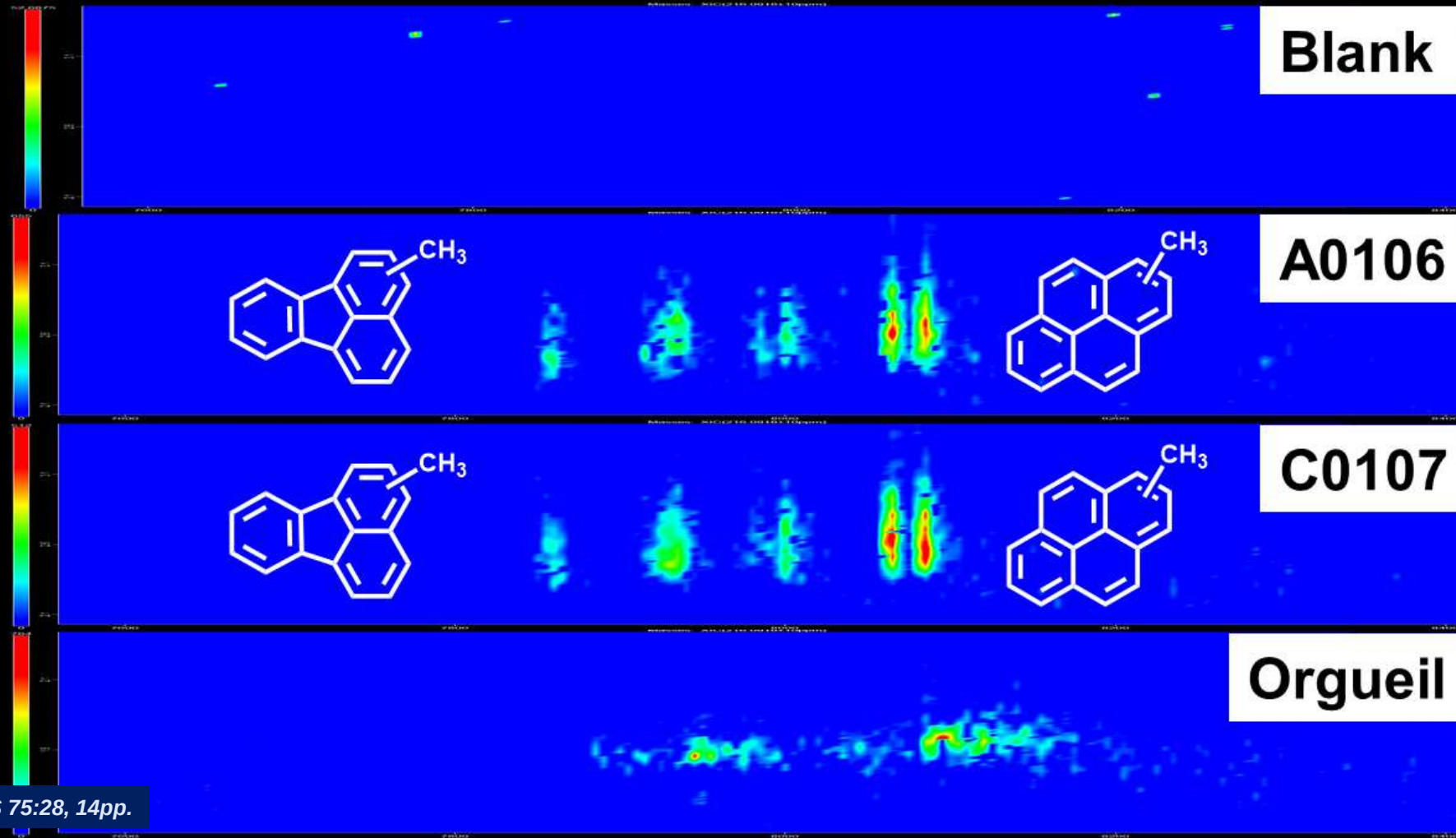
Ryugu: Fluoranthene and Pyrene

Fluoranthene and pyrene, m/z 202.0778 $\pm$ 10 ppm



Ryugu: Methylfluoranthene and Methylpyrene

Methylfluoranthene and methylpyrene,
 m/z 216.0918 $\pm$ 10 ppm



Reaction Pathways for PAHs

